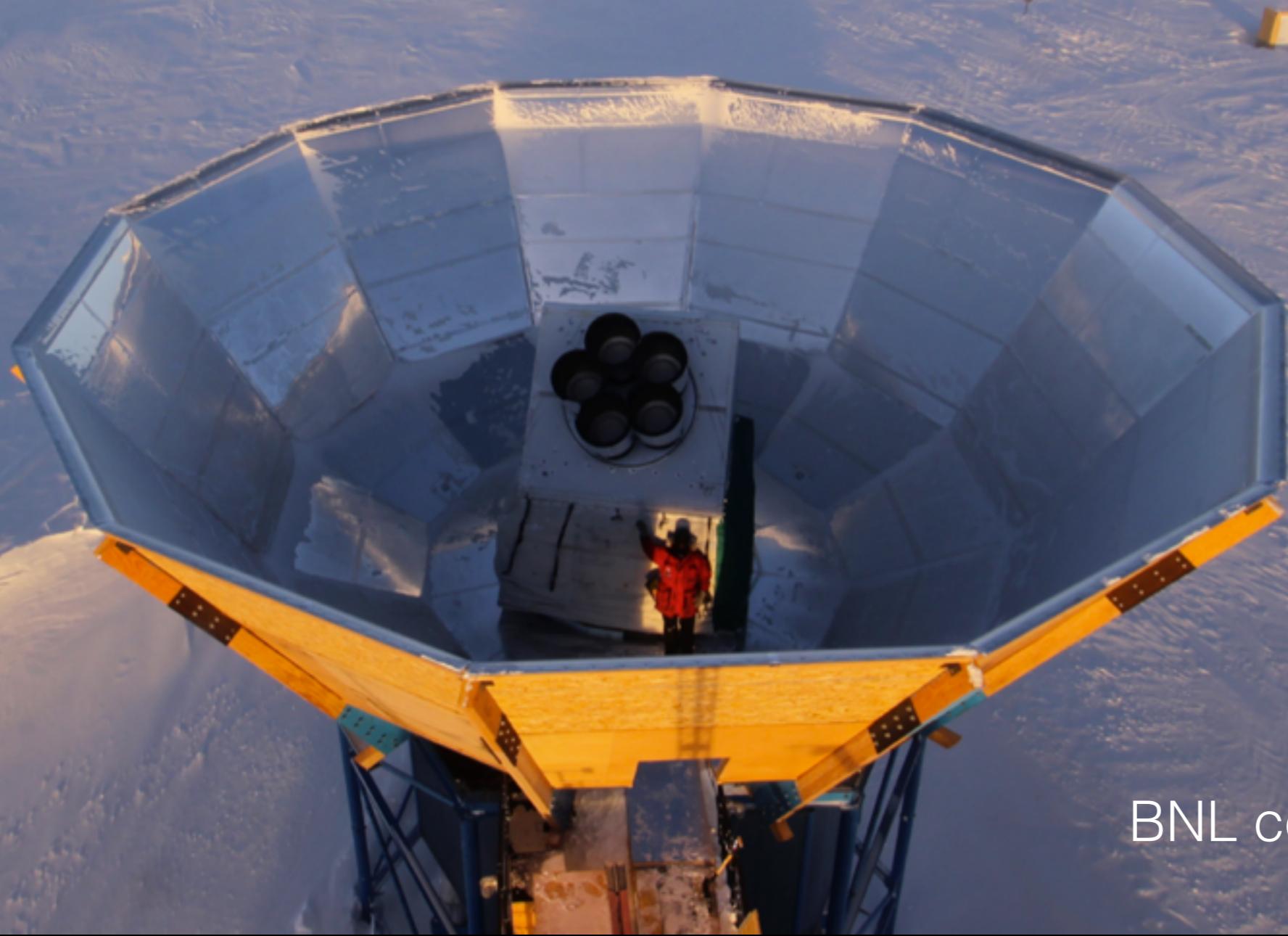


Update from the BICEP/Keck Program



Chris Sheehy
BNL cosmo group meeting
Feb 20, 2017

Ground based CMB at the South Pole

South Pole CMB telescopes



10m South Pole Telescope

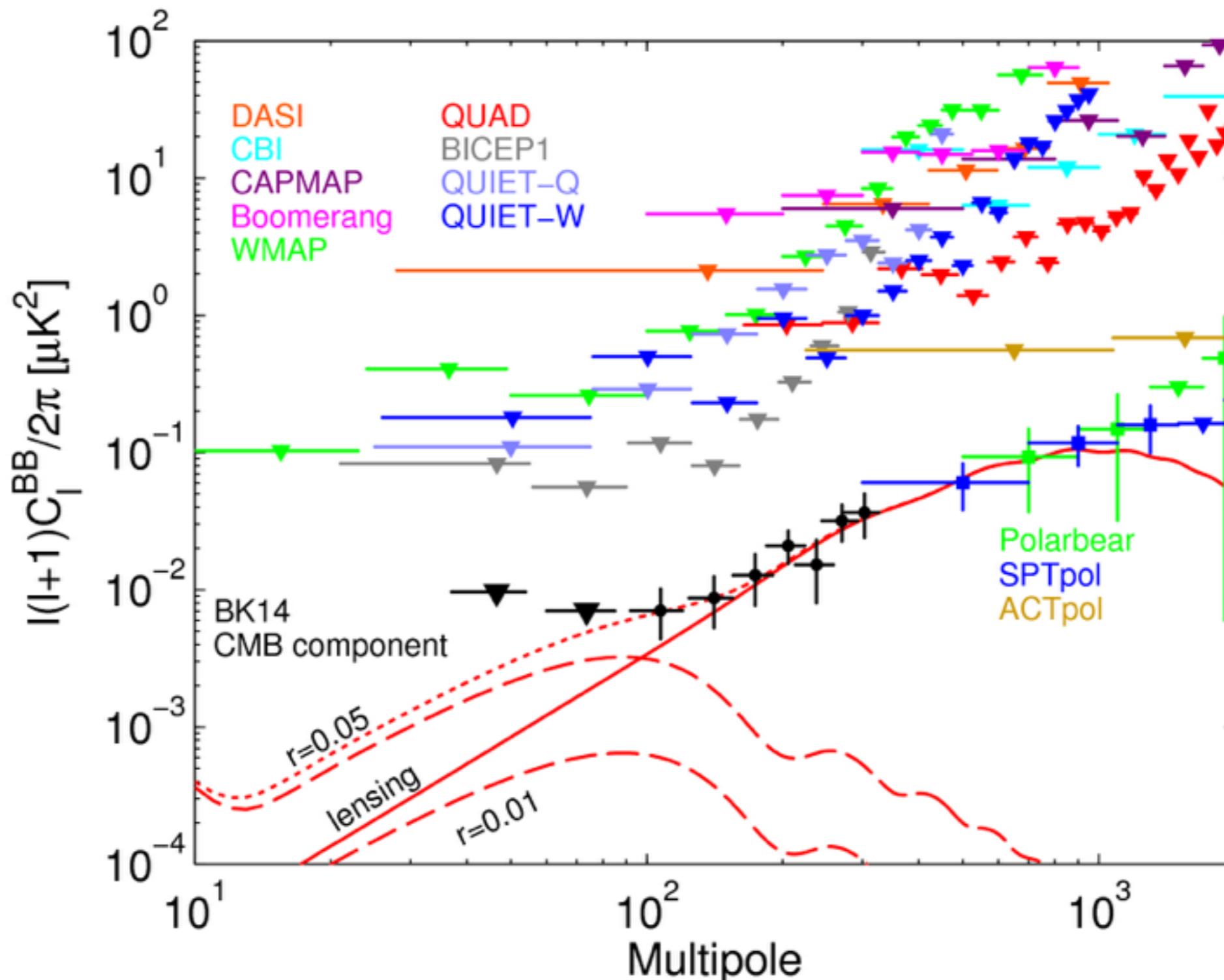
BICEP1
BICEP2
BICEP3

SPIDER

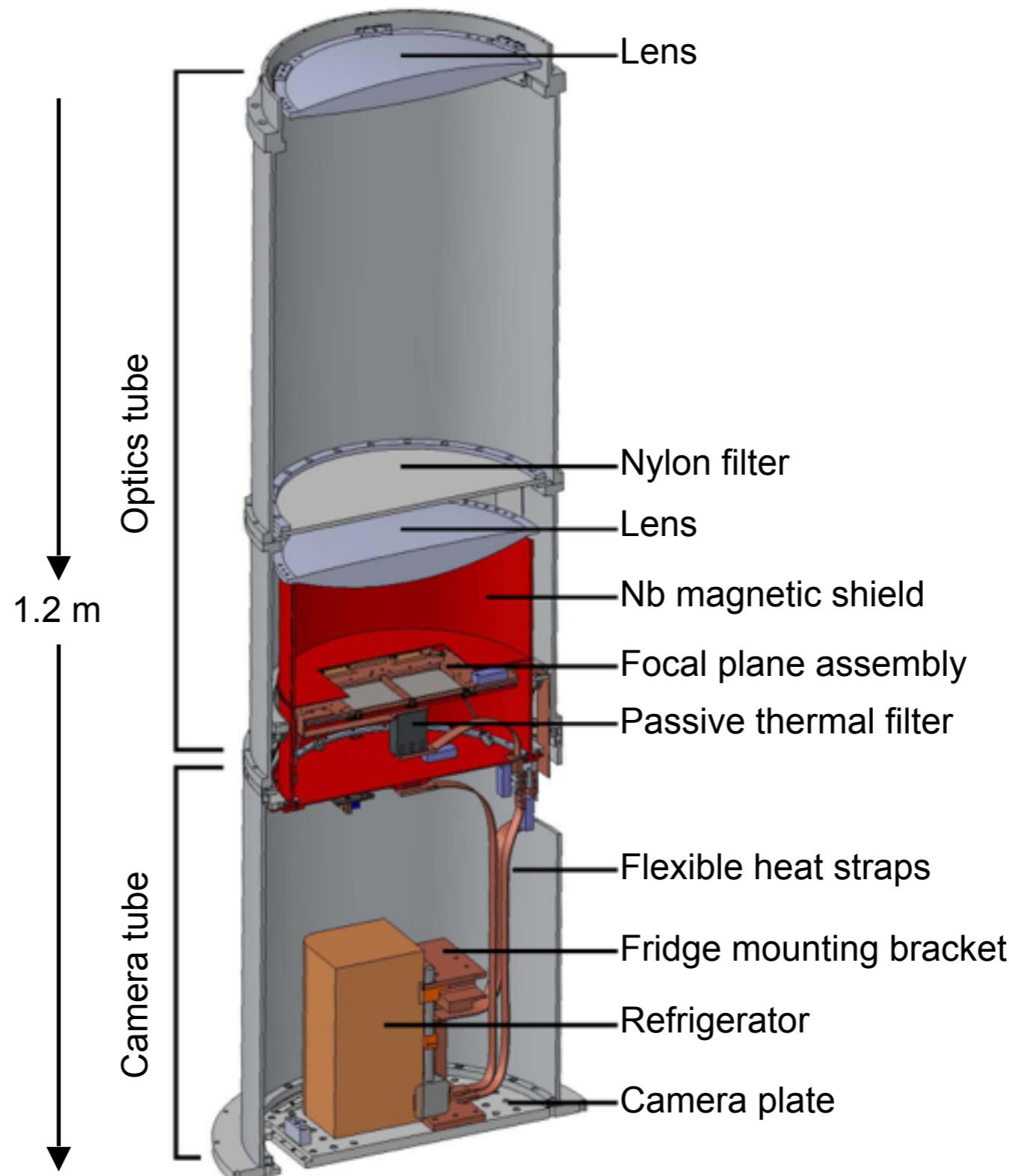


DASI
QUAD
Keck
Array

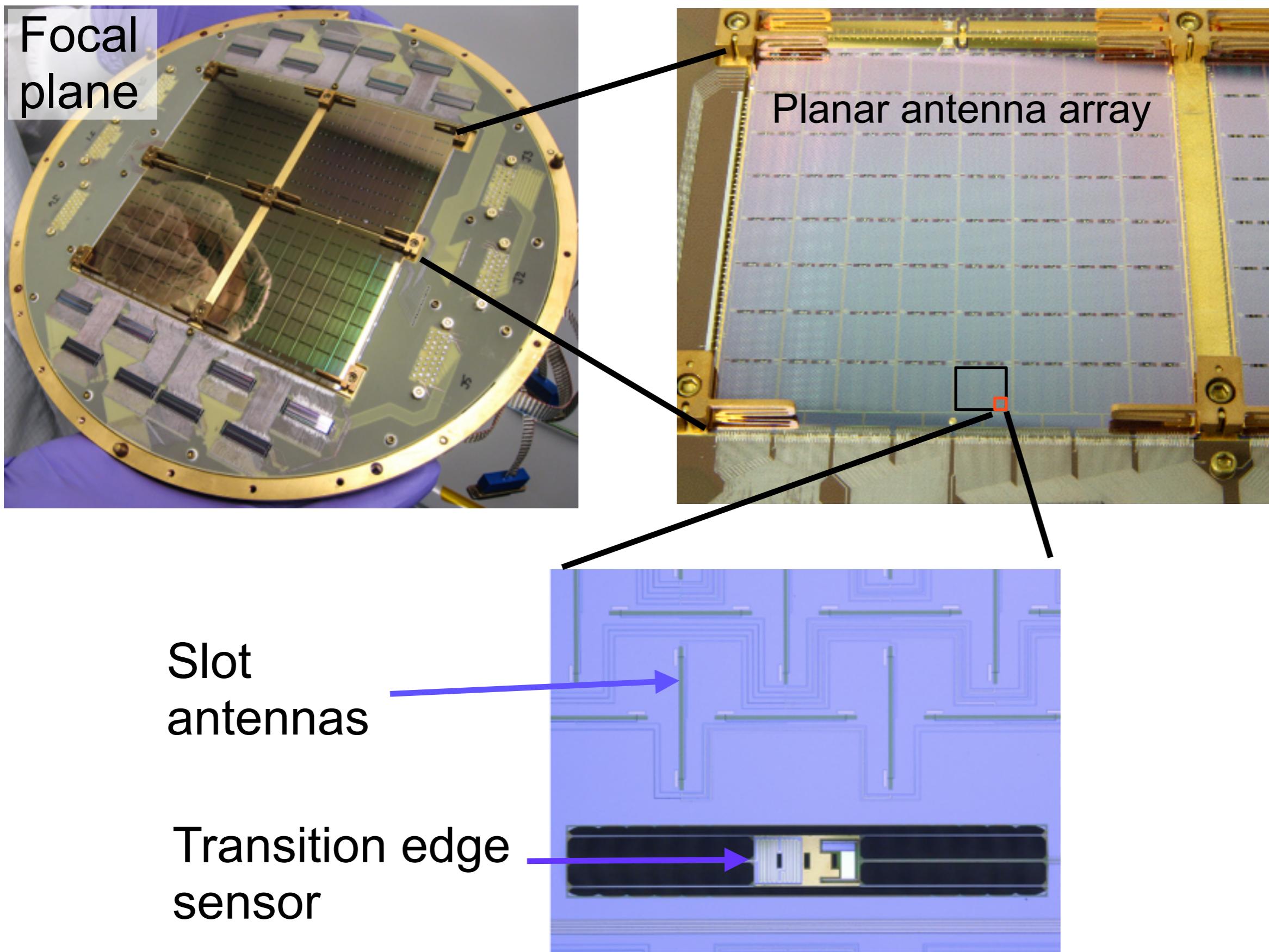
Current state of CMB B-modes



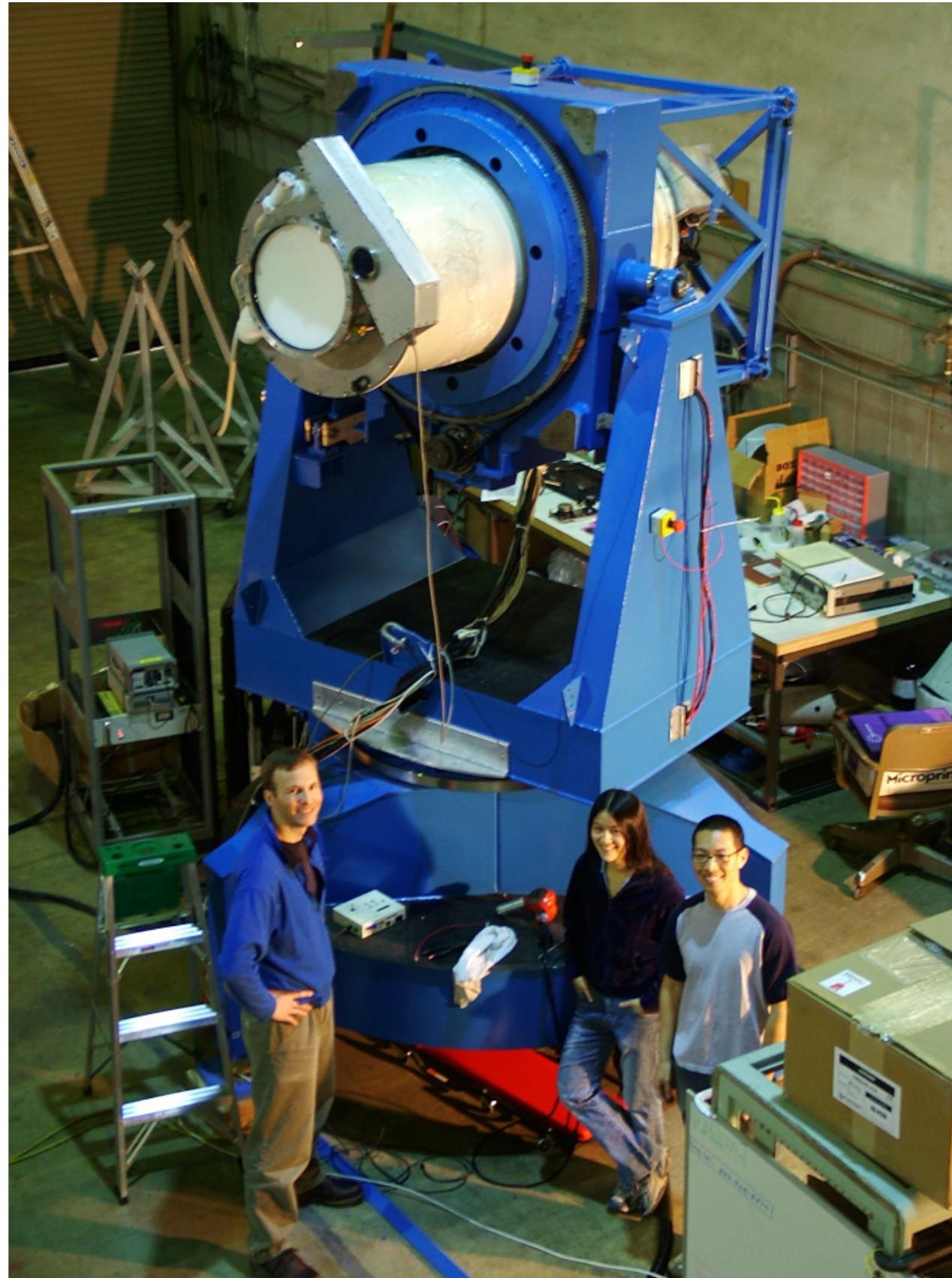
The BICEP2/Keck Telescopes



Mass-produced Superconducting Detectors



The BICEP2 Telescope

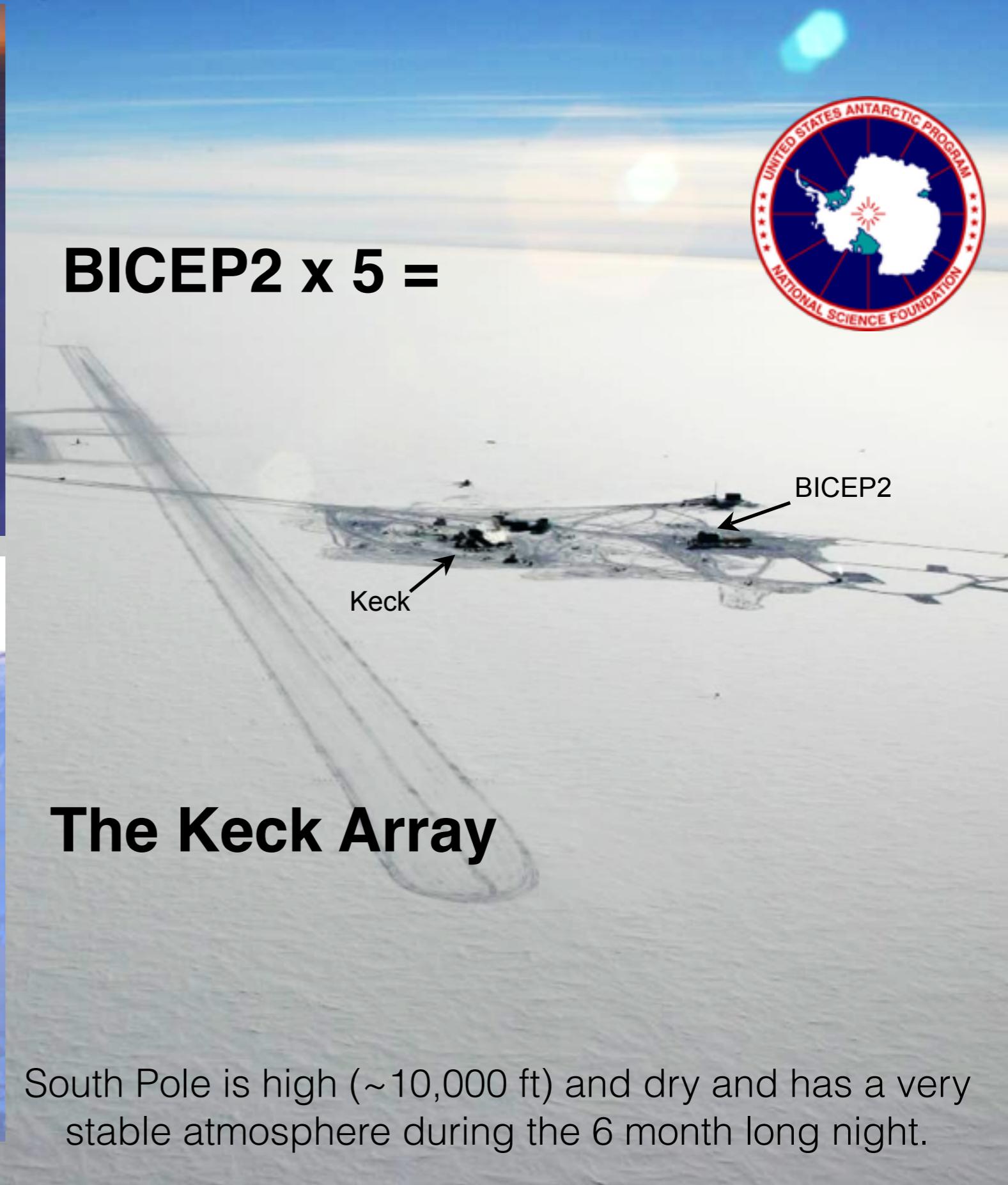


BICEP1

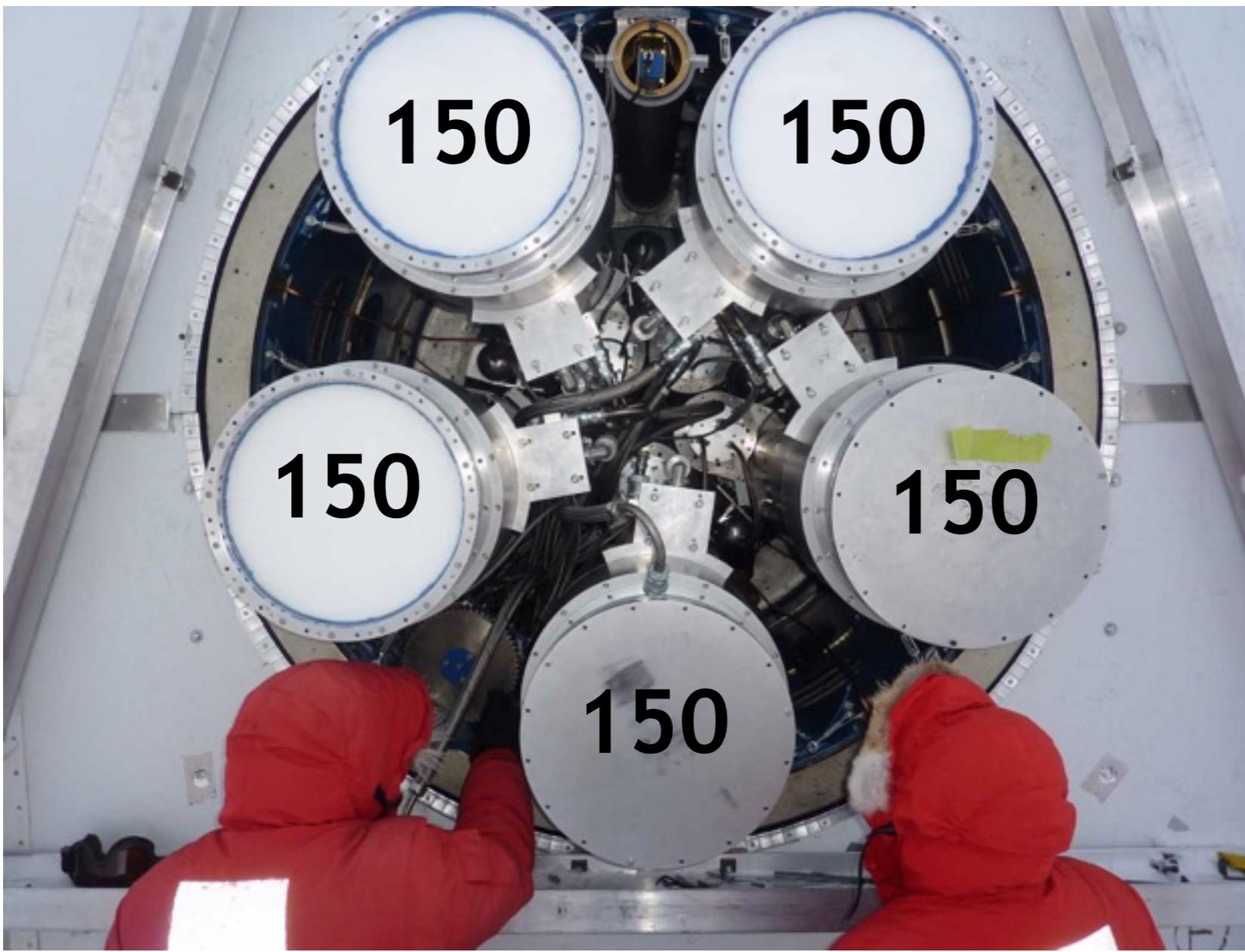
The Keck Telescopes



BICEP2 and Keck Array

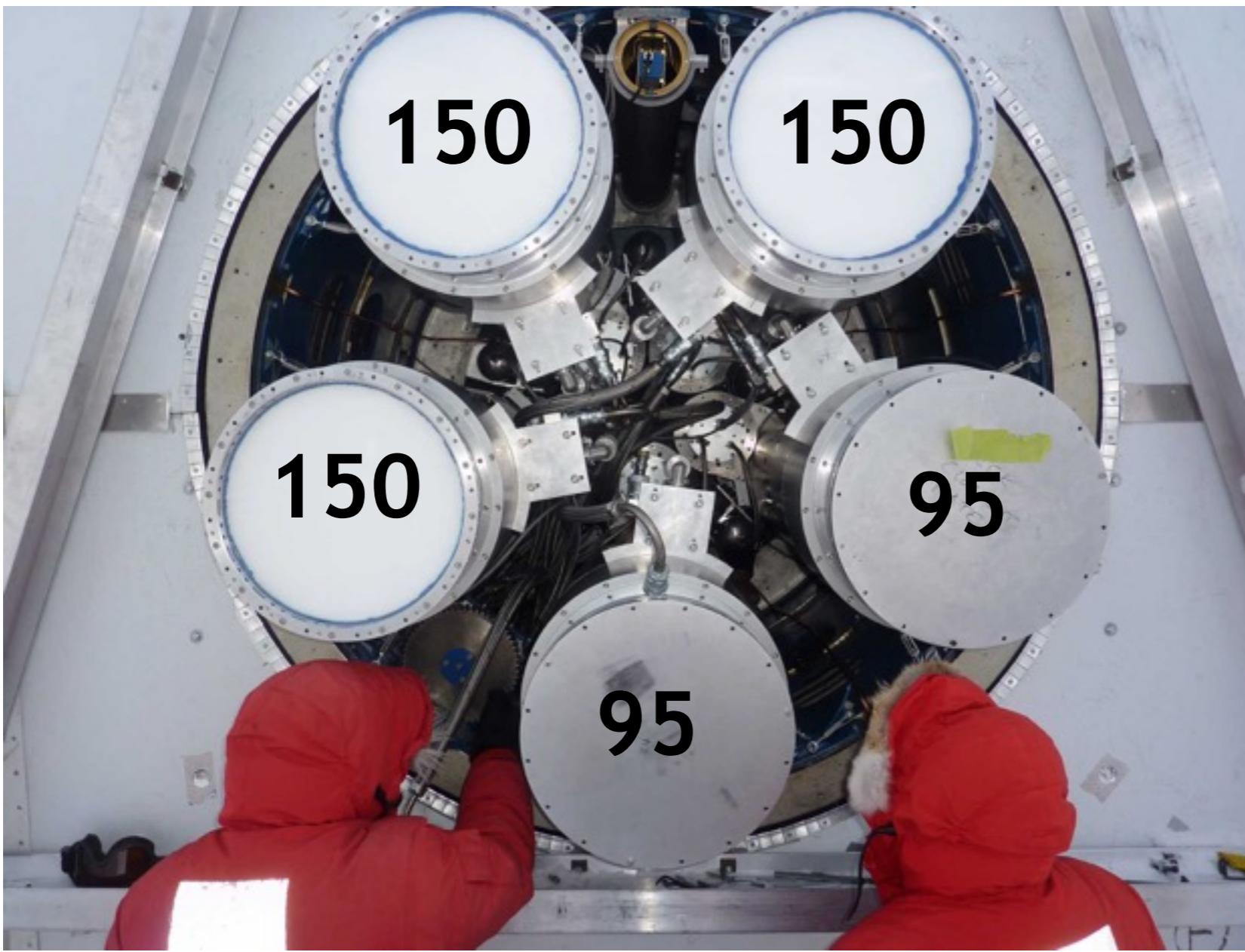


Keck Array Frequency Coverage



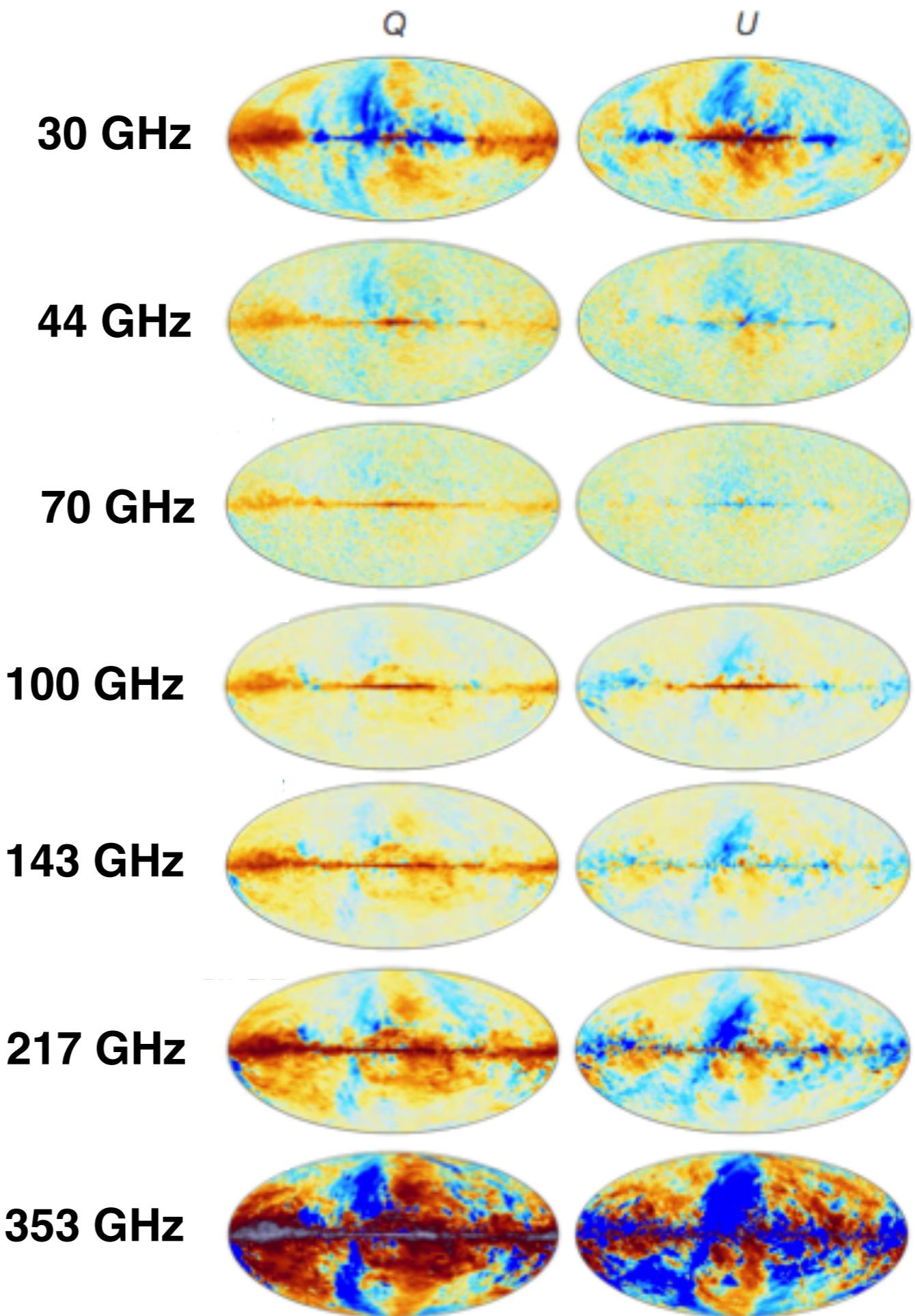
2012-2013

Keck Array Frequency Coverage

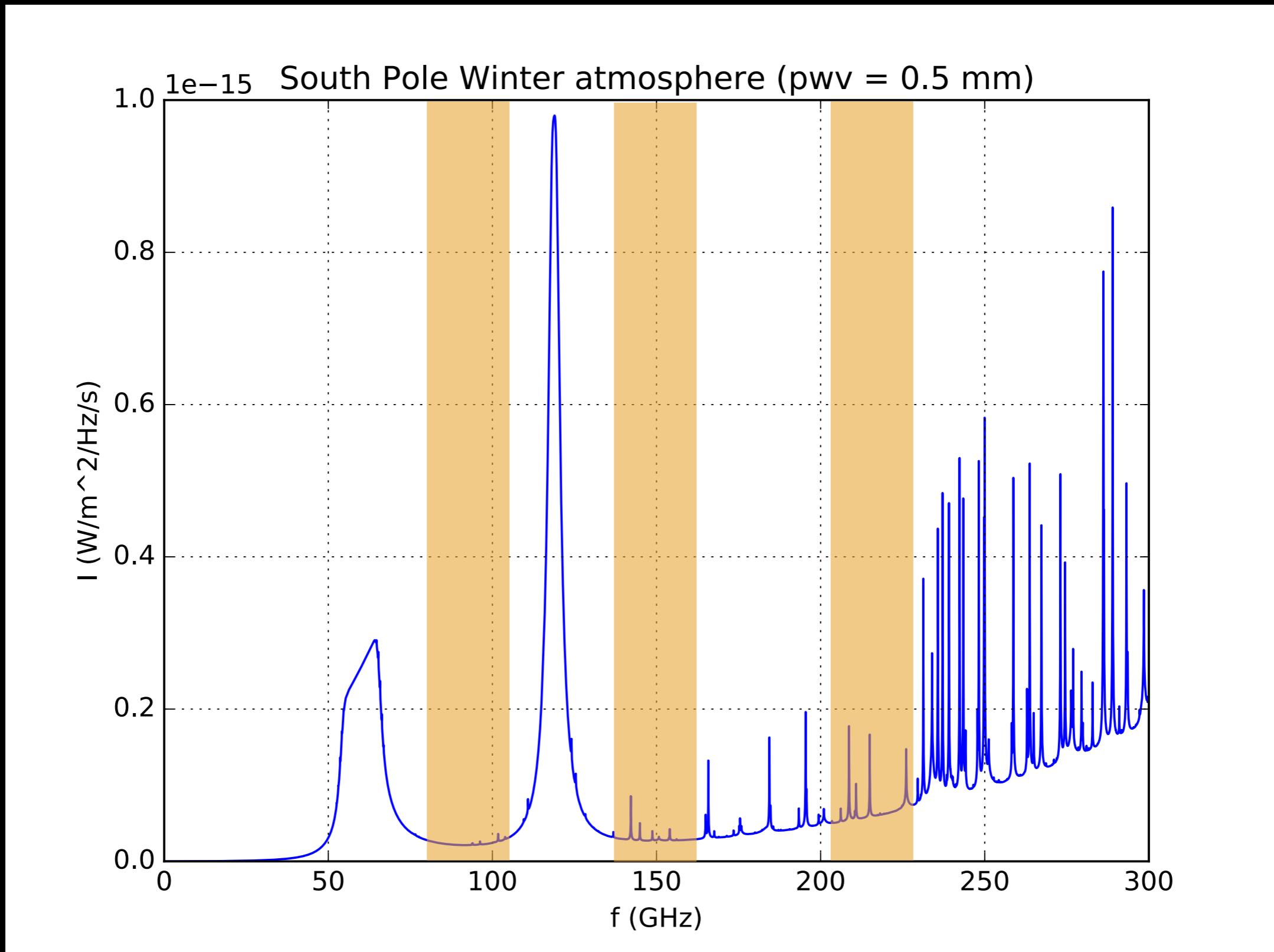


2014

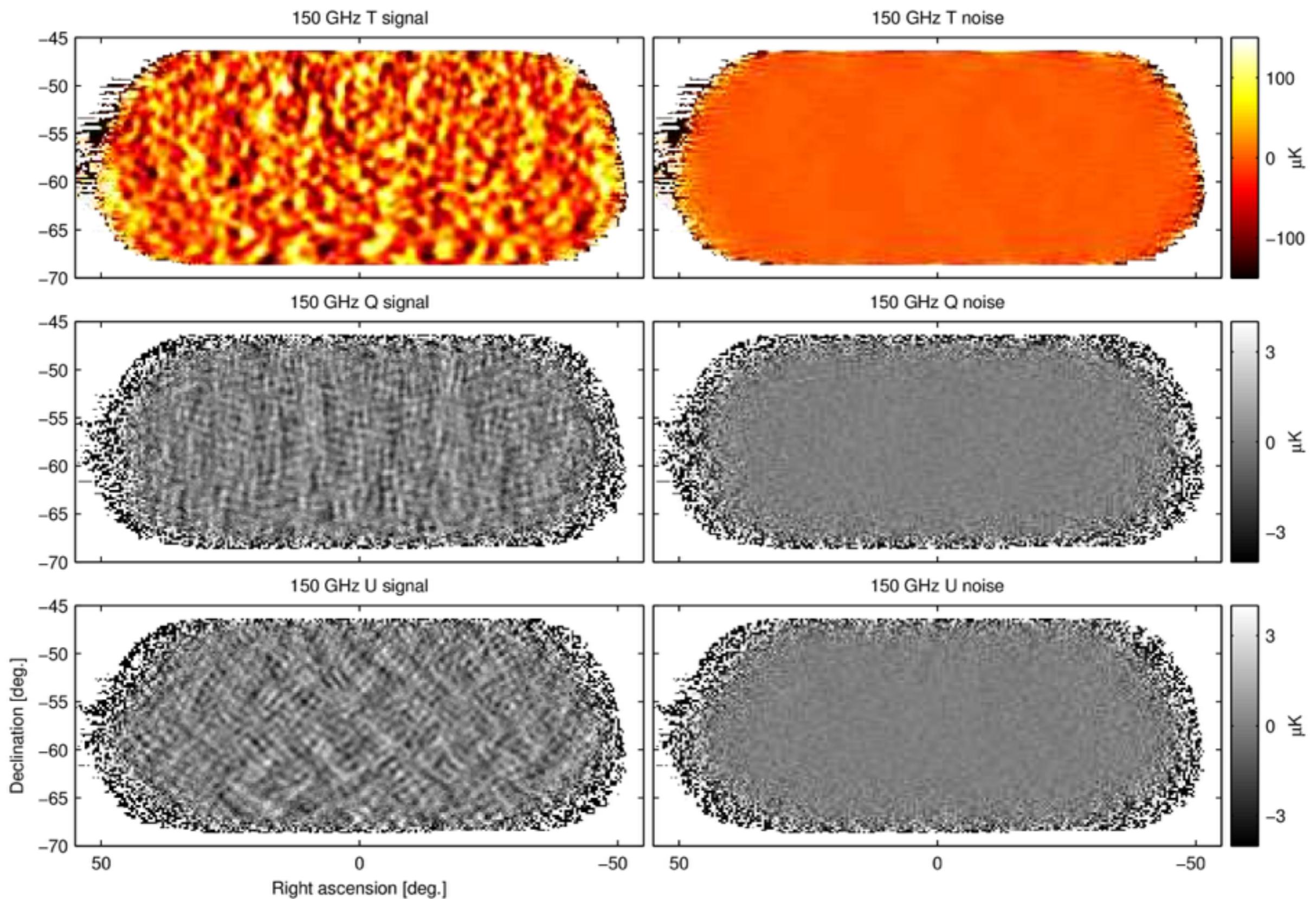
Galactic foregrounds:
synchrotron at low
frequencies, thermally
emitting dust (~ 20 K) at
high frequencies



Ground windows

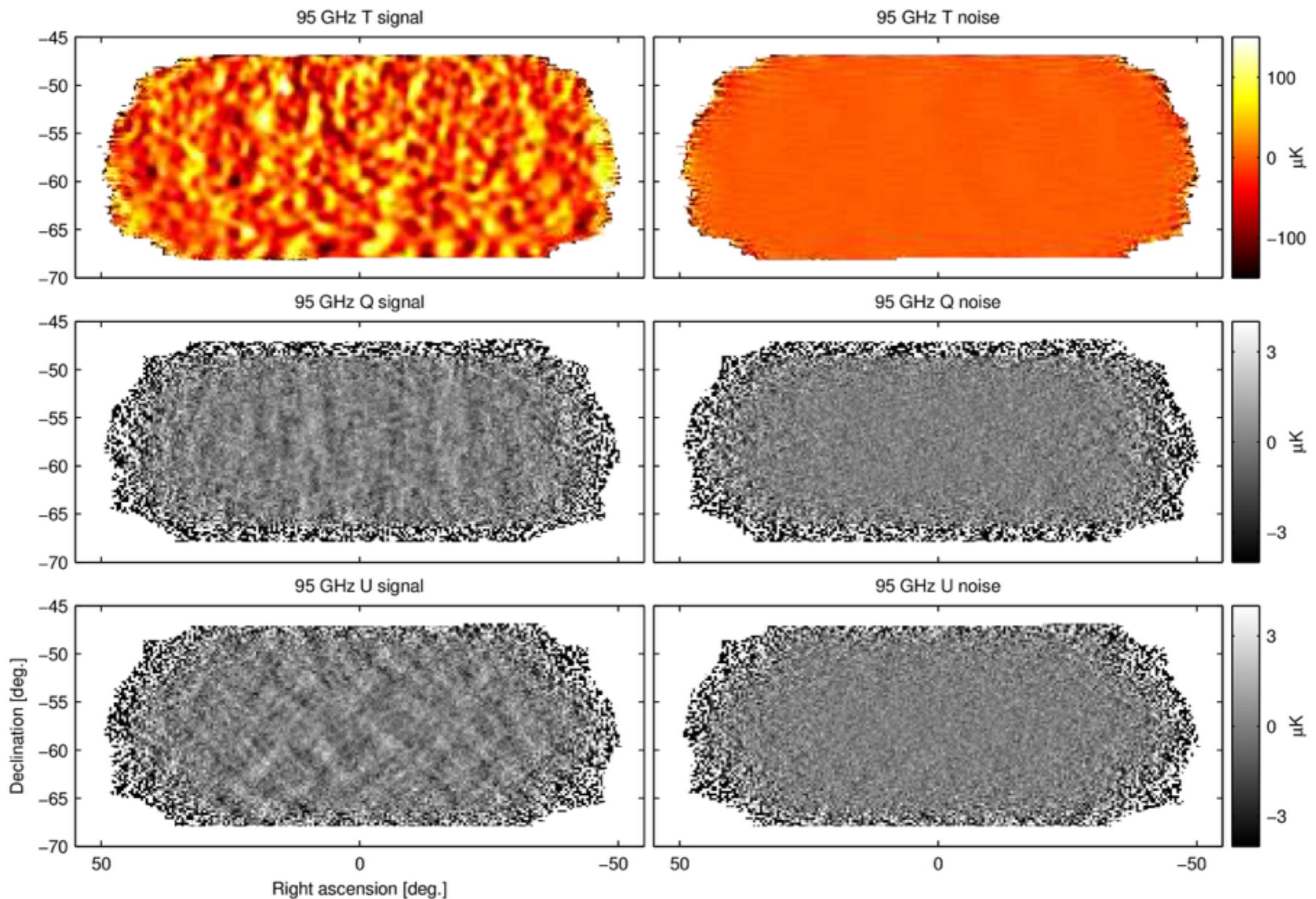


150 GHz maps

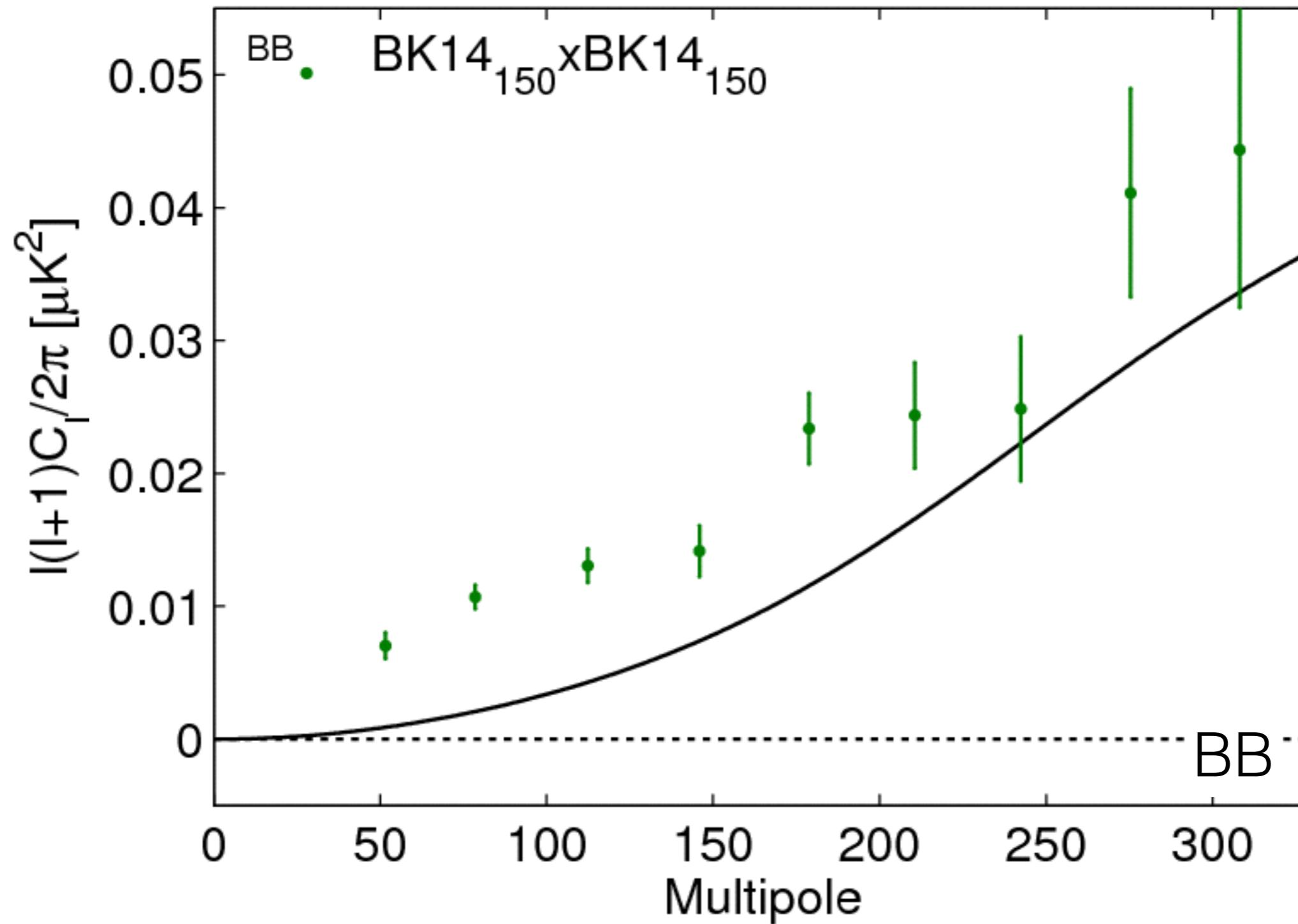


BK14 150GHz – 50 nK deg (3.0 μK arcmin)

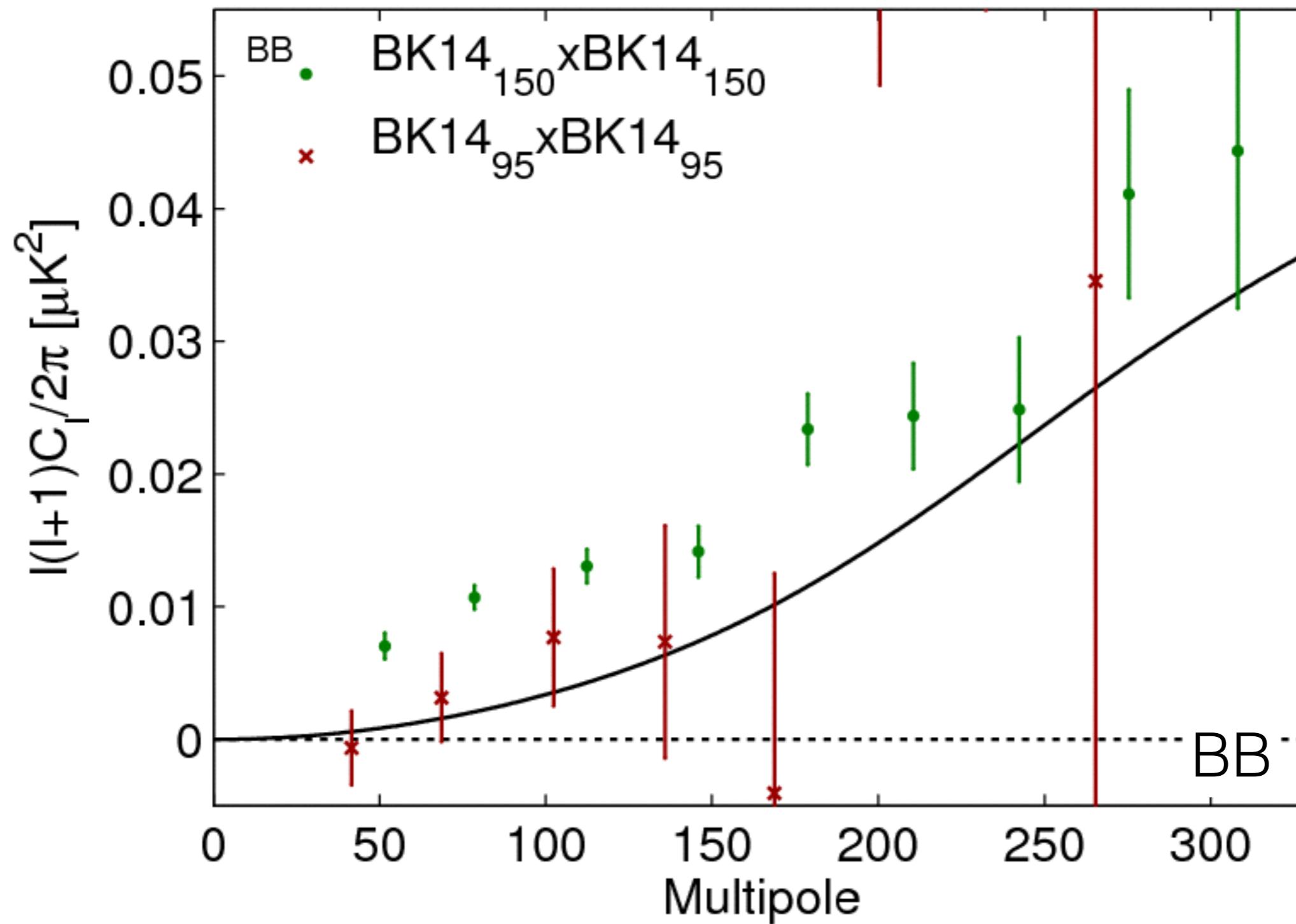
95 GHz maps



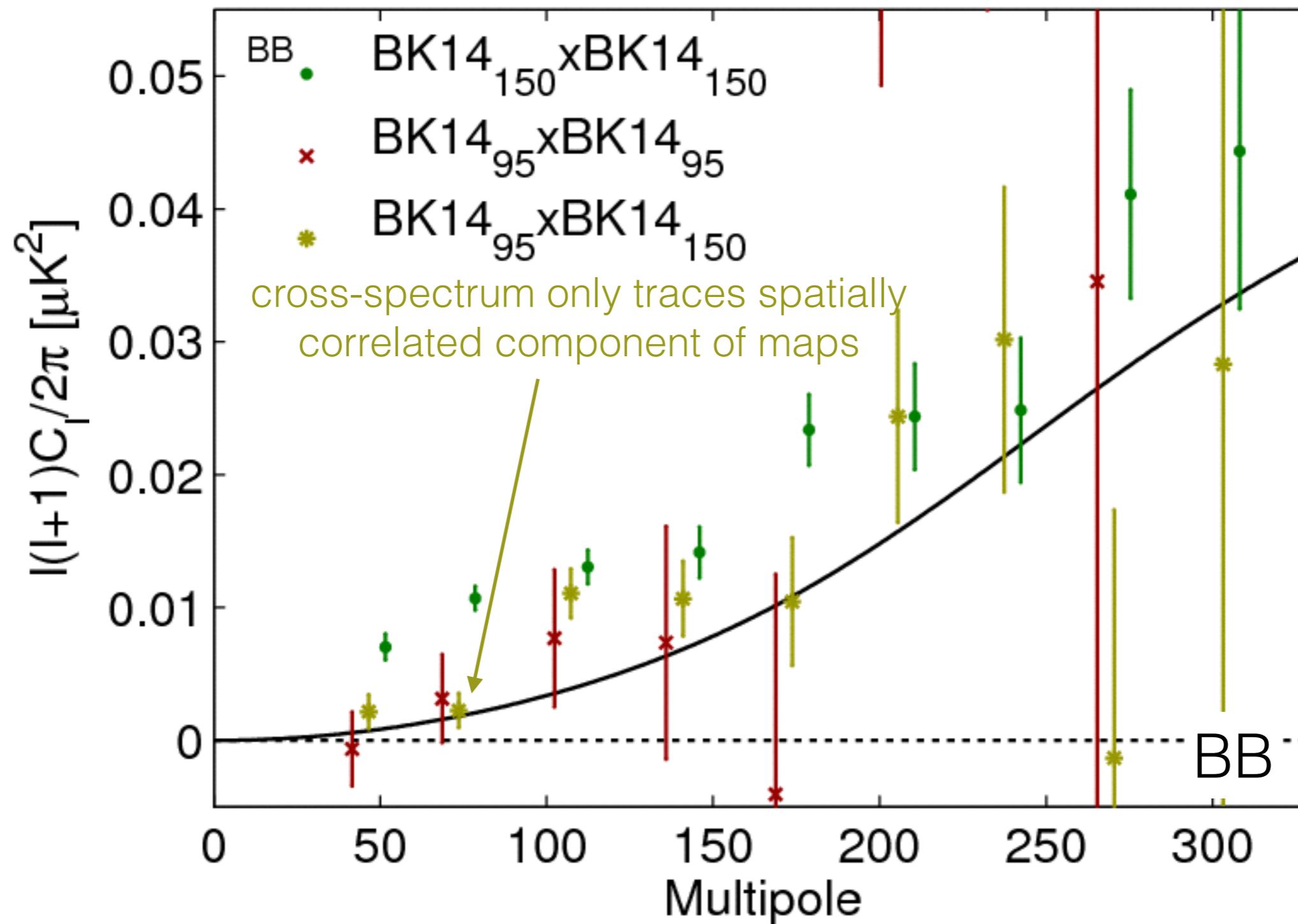
BICEP2 + Keck BB auto and cross-spectra



BICEP2 + Keck BB auto and cross-spectra



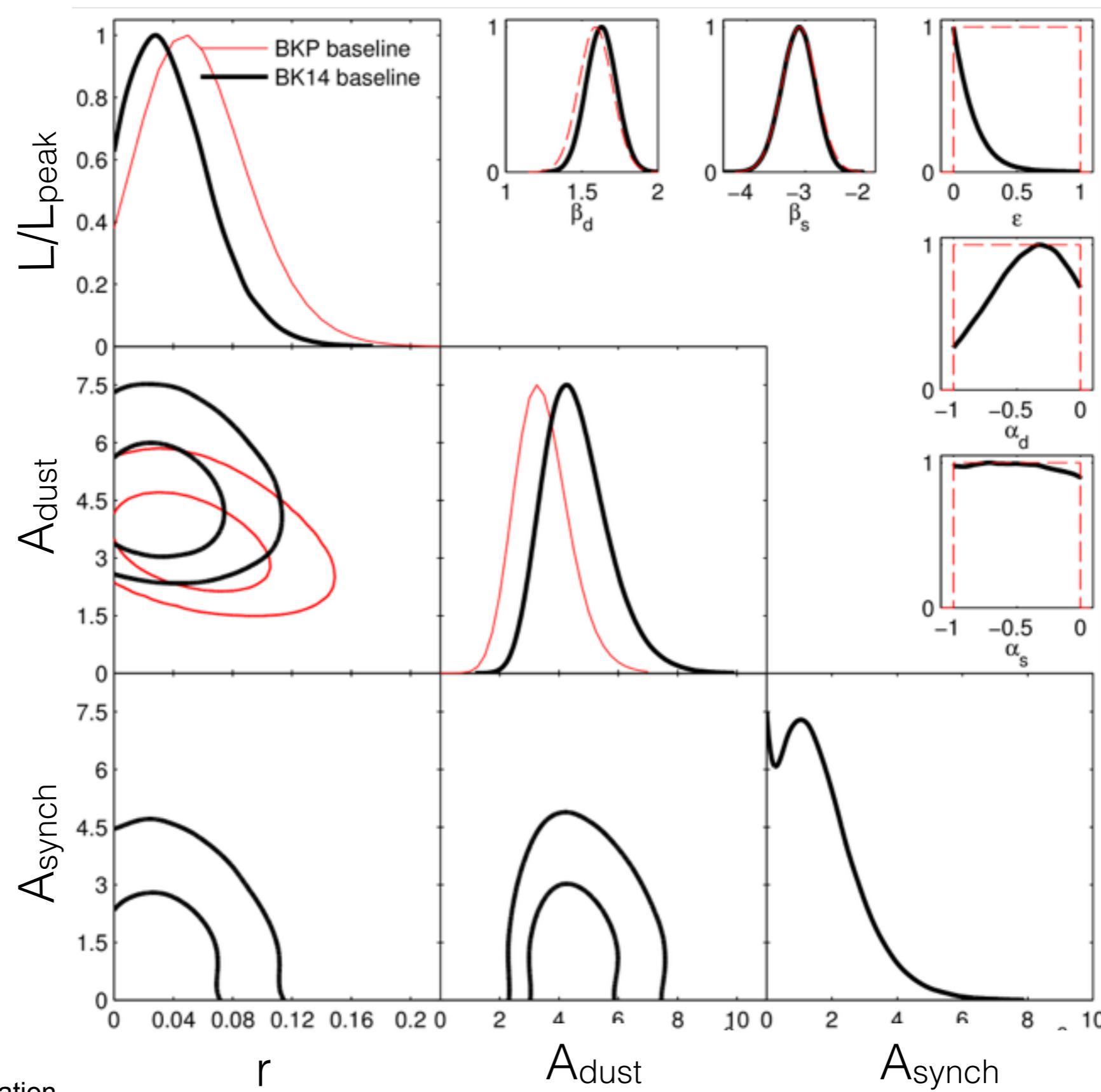
BICEP2 + Keck BB auto and cross-spectra



Put priors on the frequency spectral
indices of dust & sync

Allow dust/sync
correlation

dust vs. r →
degeneracy lifted



Put priors on the frequency spectral indices of dust & sync

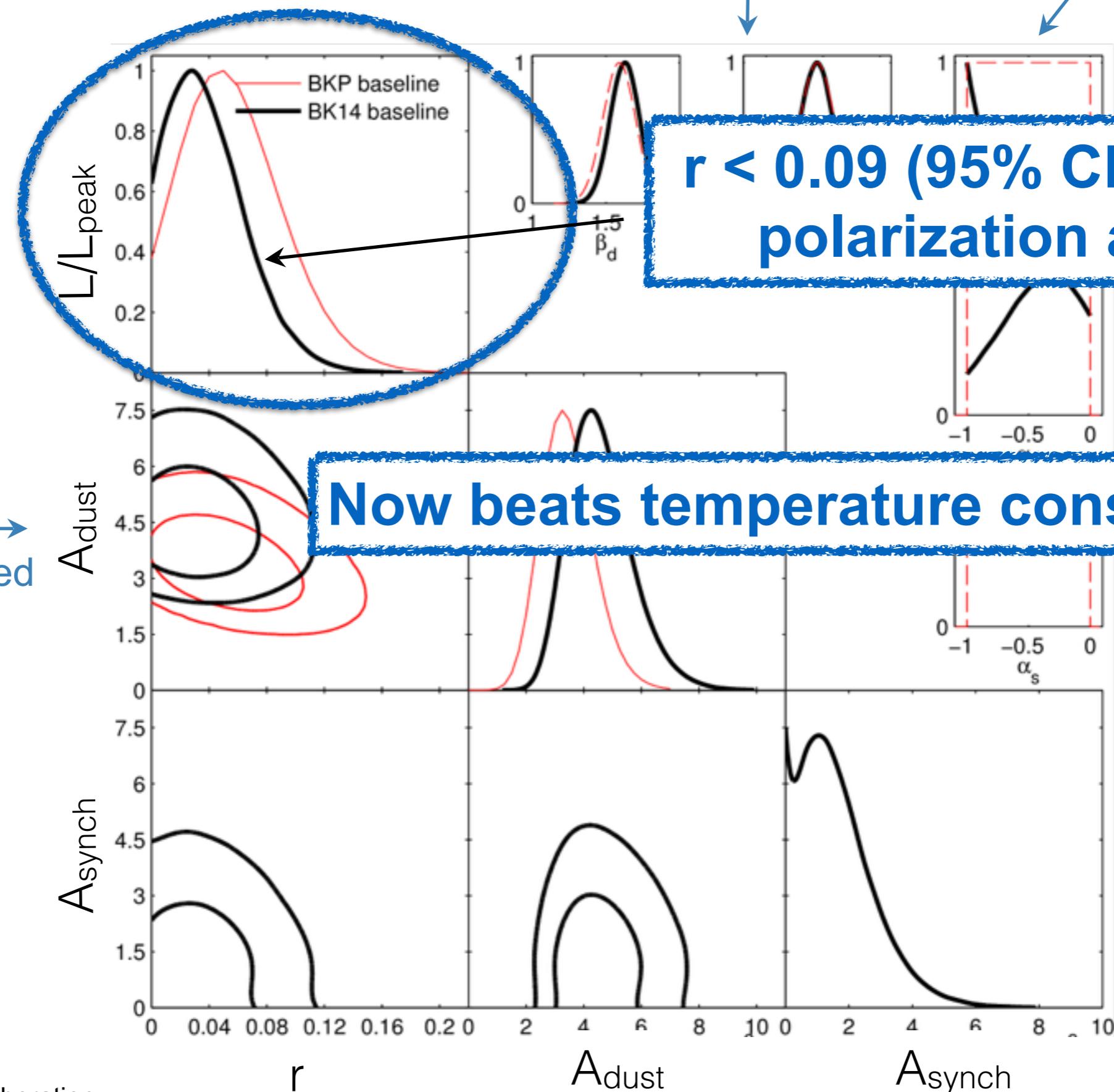
Allow dust/sync correlation

$r < 0.09$ (95% CL) using polarization alone

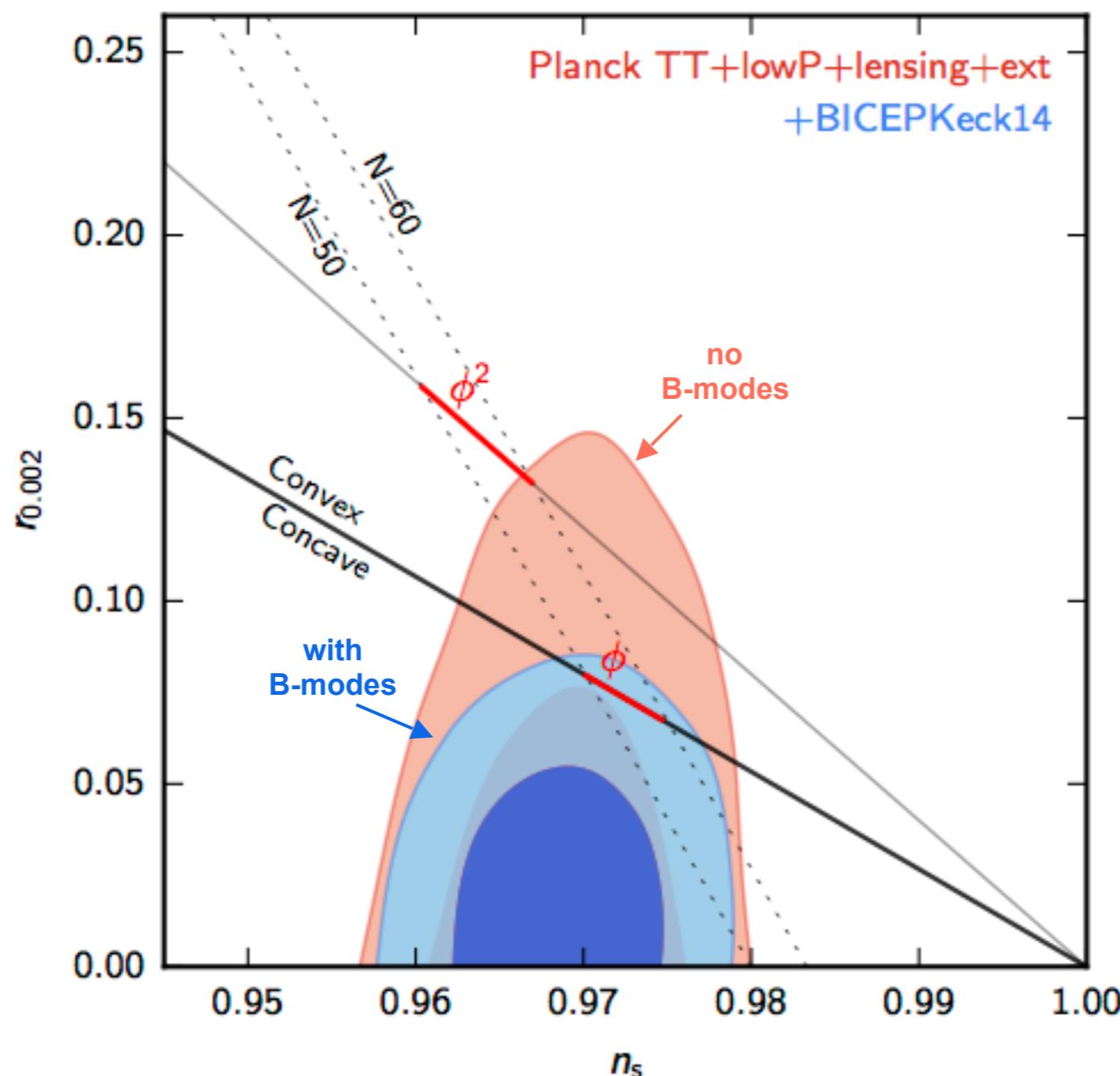
dust vs. r →
degeneracy lifted

Now beats temperature constraints

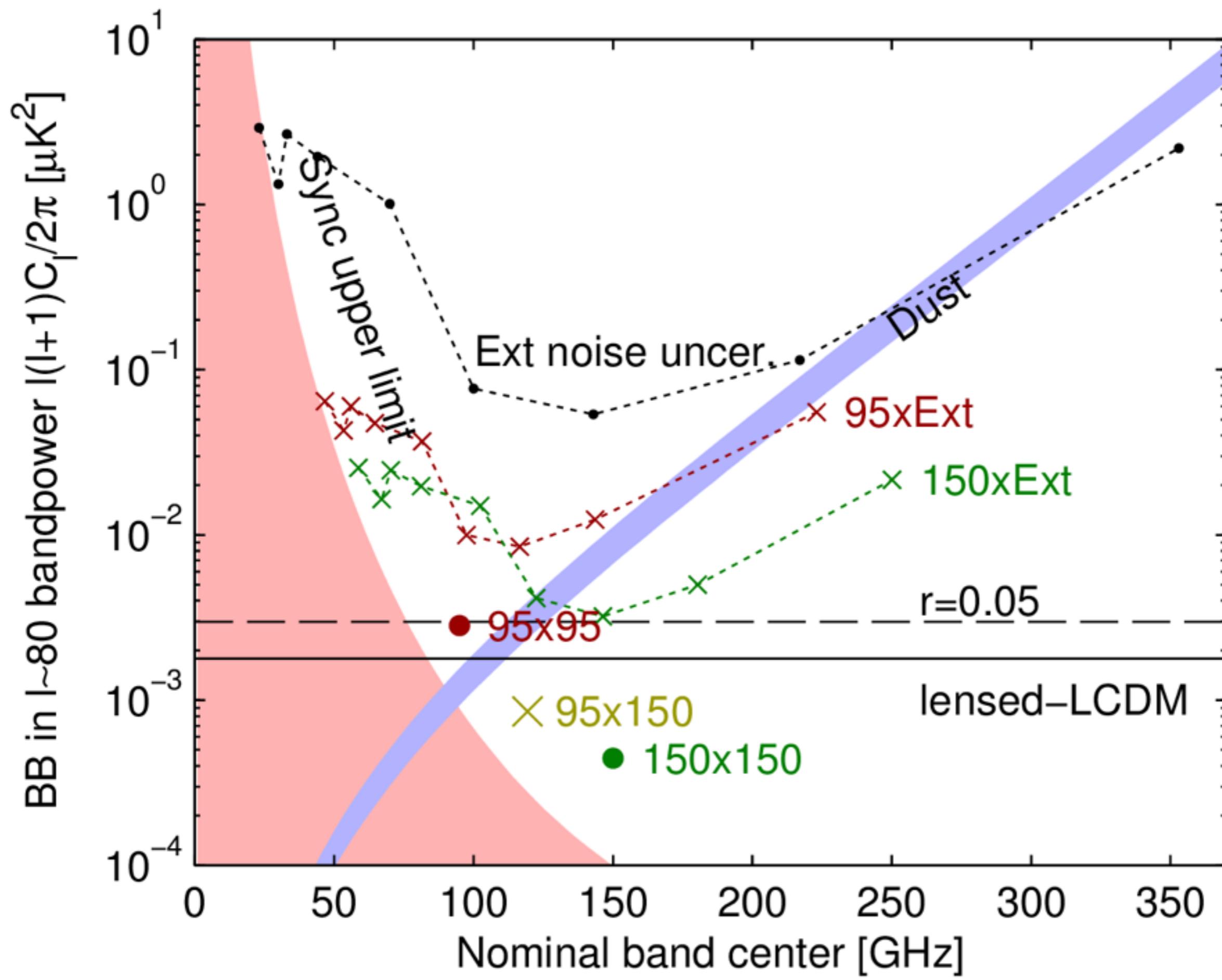
Marginalize over generous ranges in spatial spectral indices



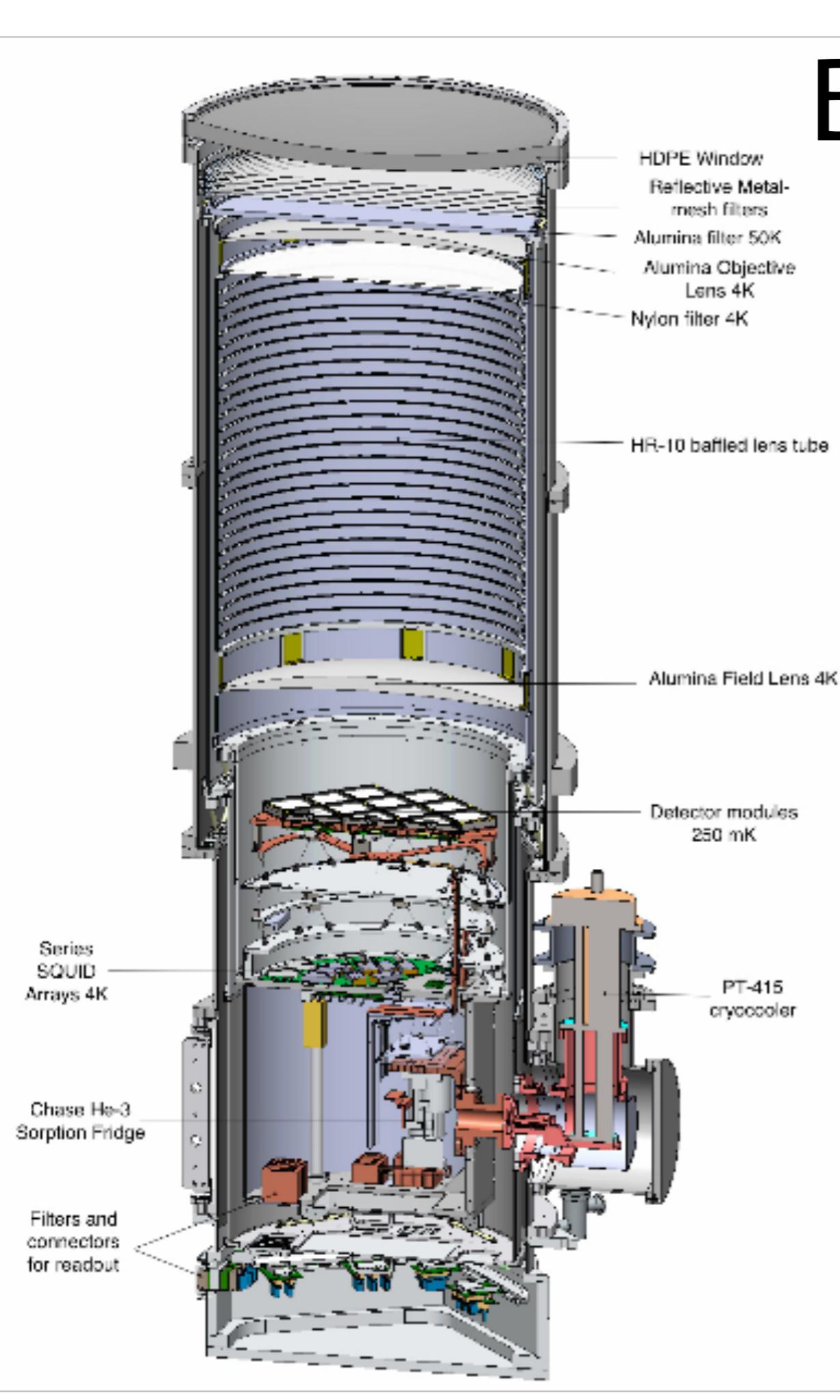
Adding in temperature



$r_{0.05} < 0.07$
(best ever)

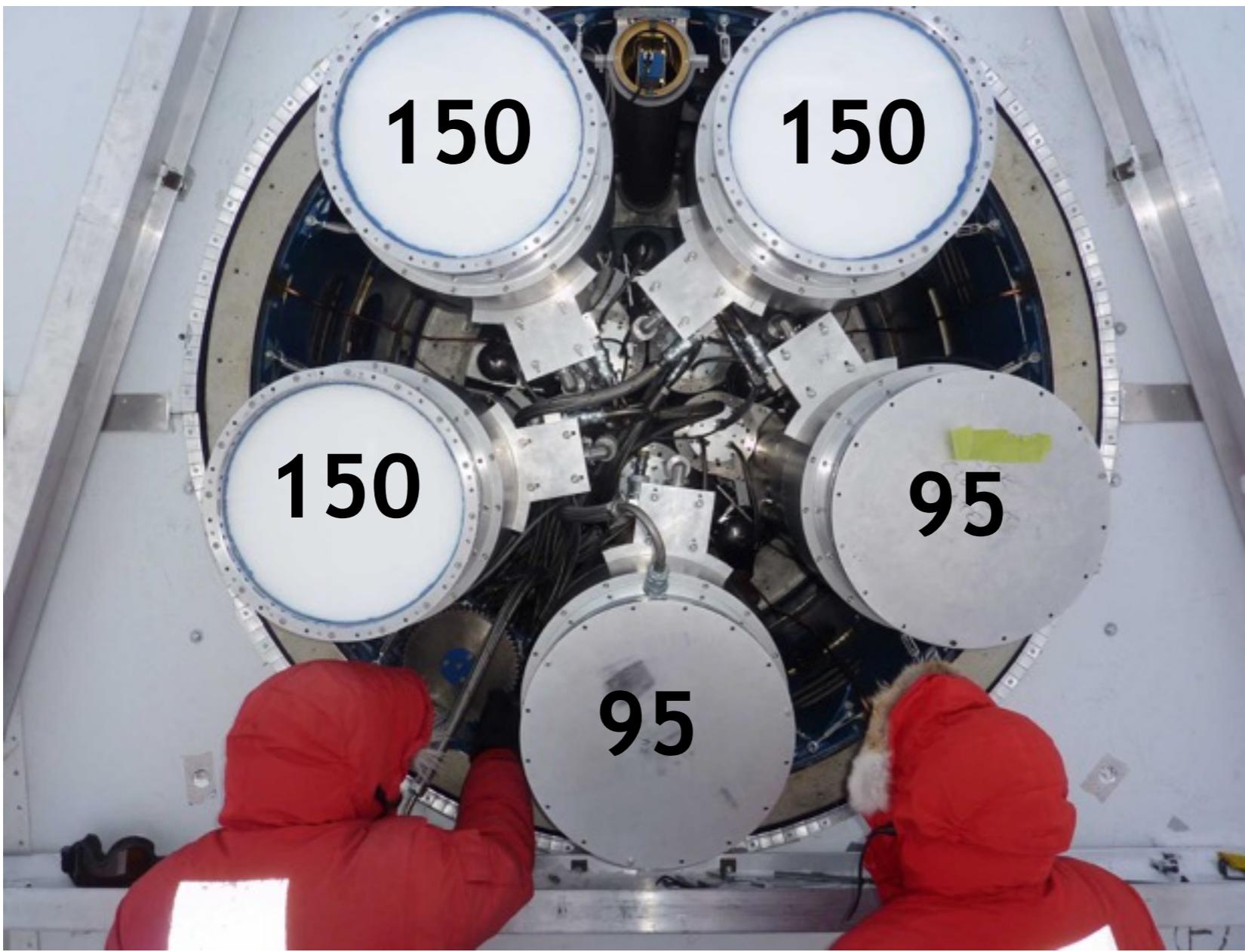


BICEP3 @ 95 GHz



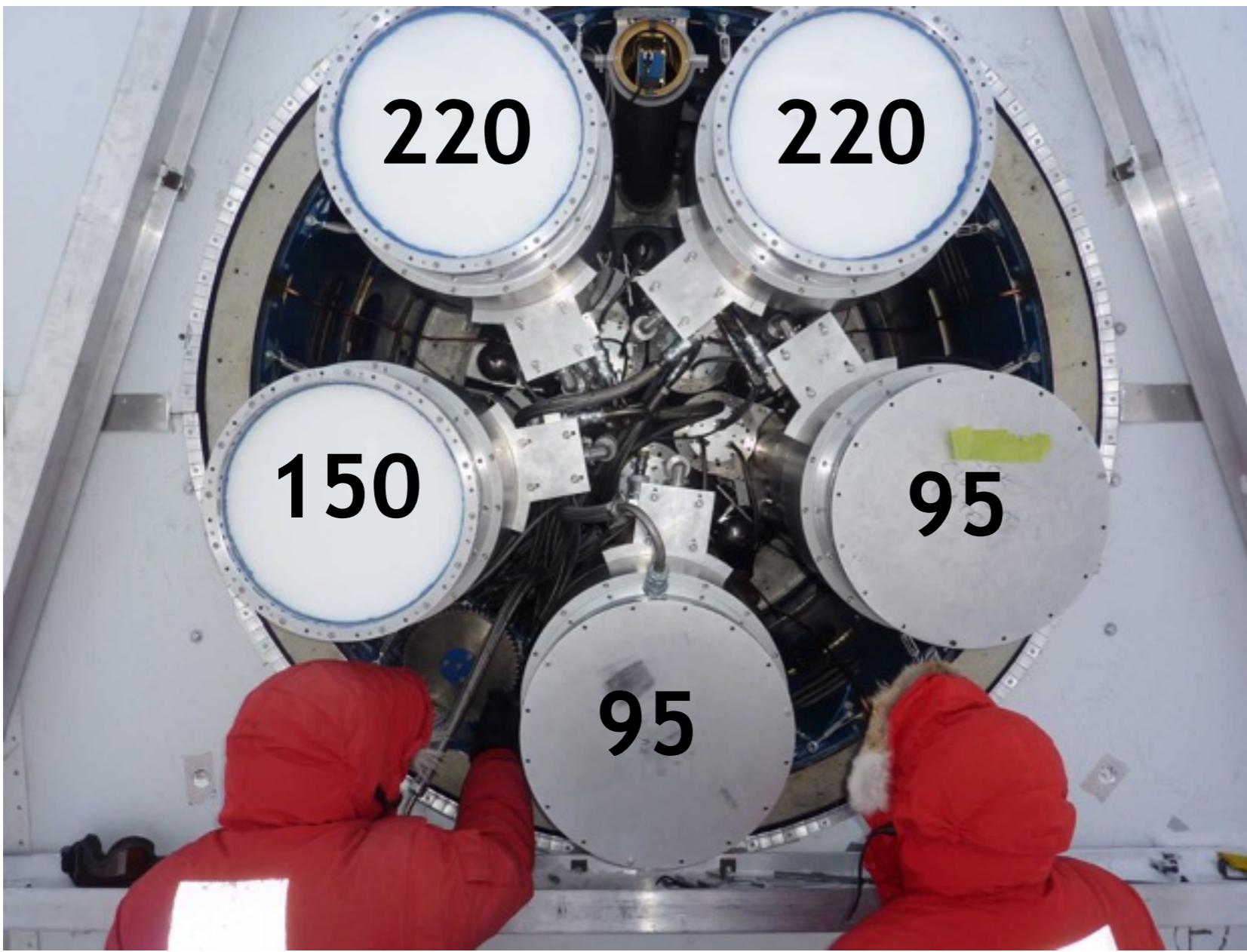
- Now taking data at the South Pole
- Roughly the same number of detectors as Keck Array, but all at 95 GHz.

Keck Array Frequency Coverage



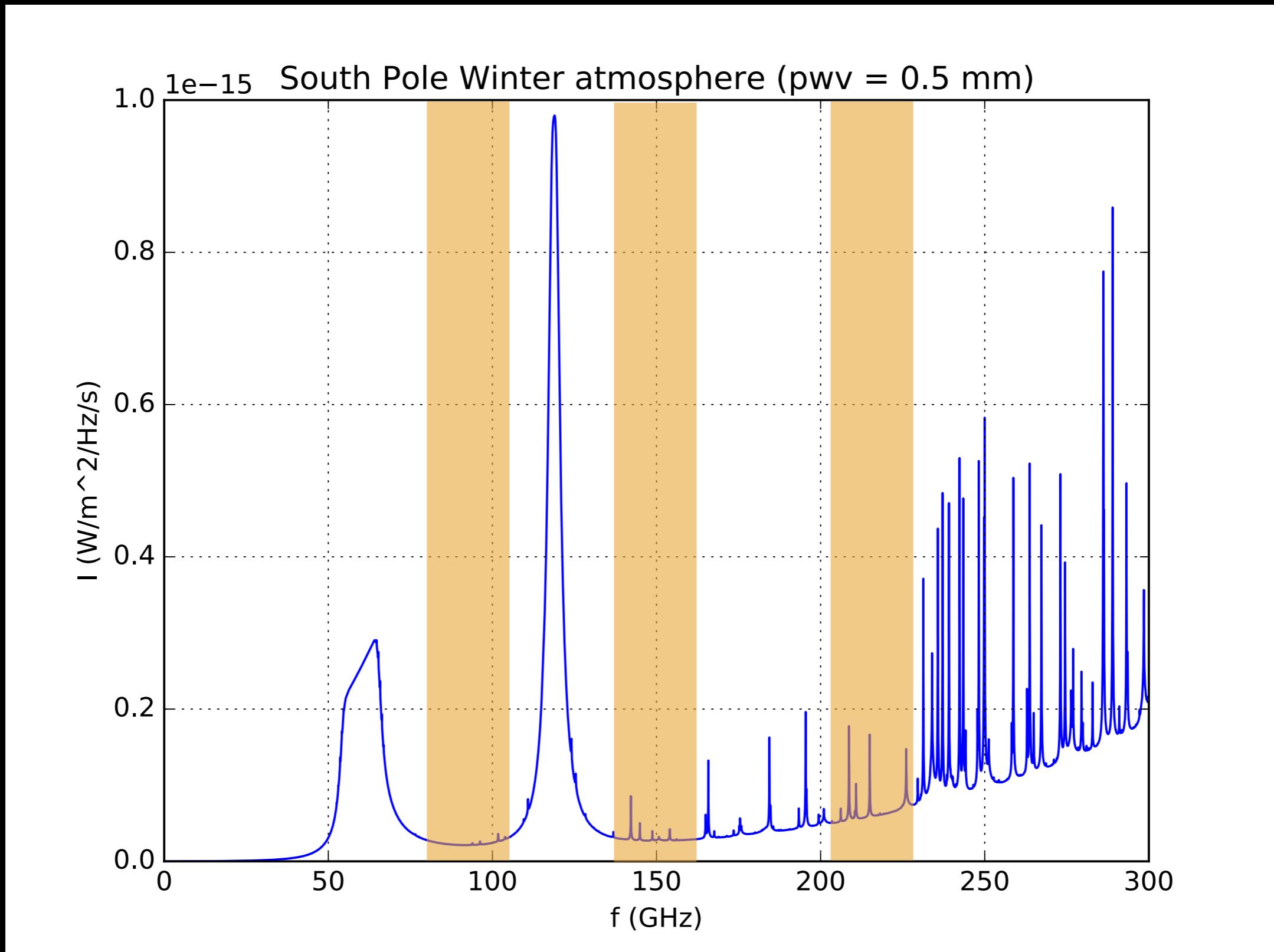
2014

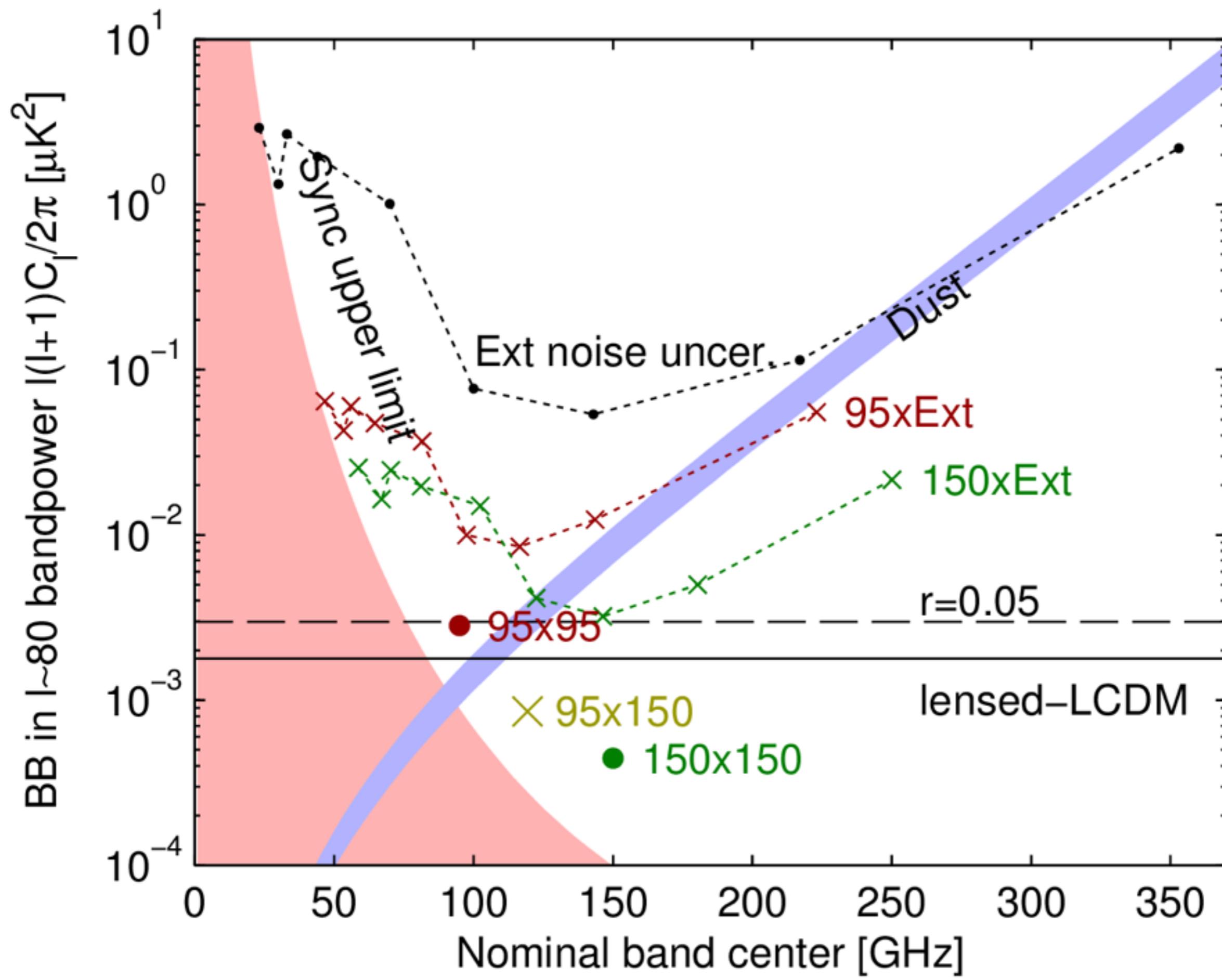
Keck Array Frequency Coverage

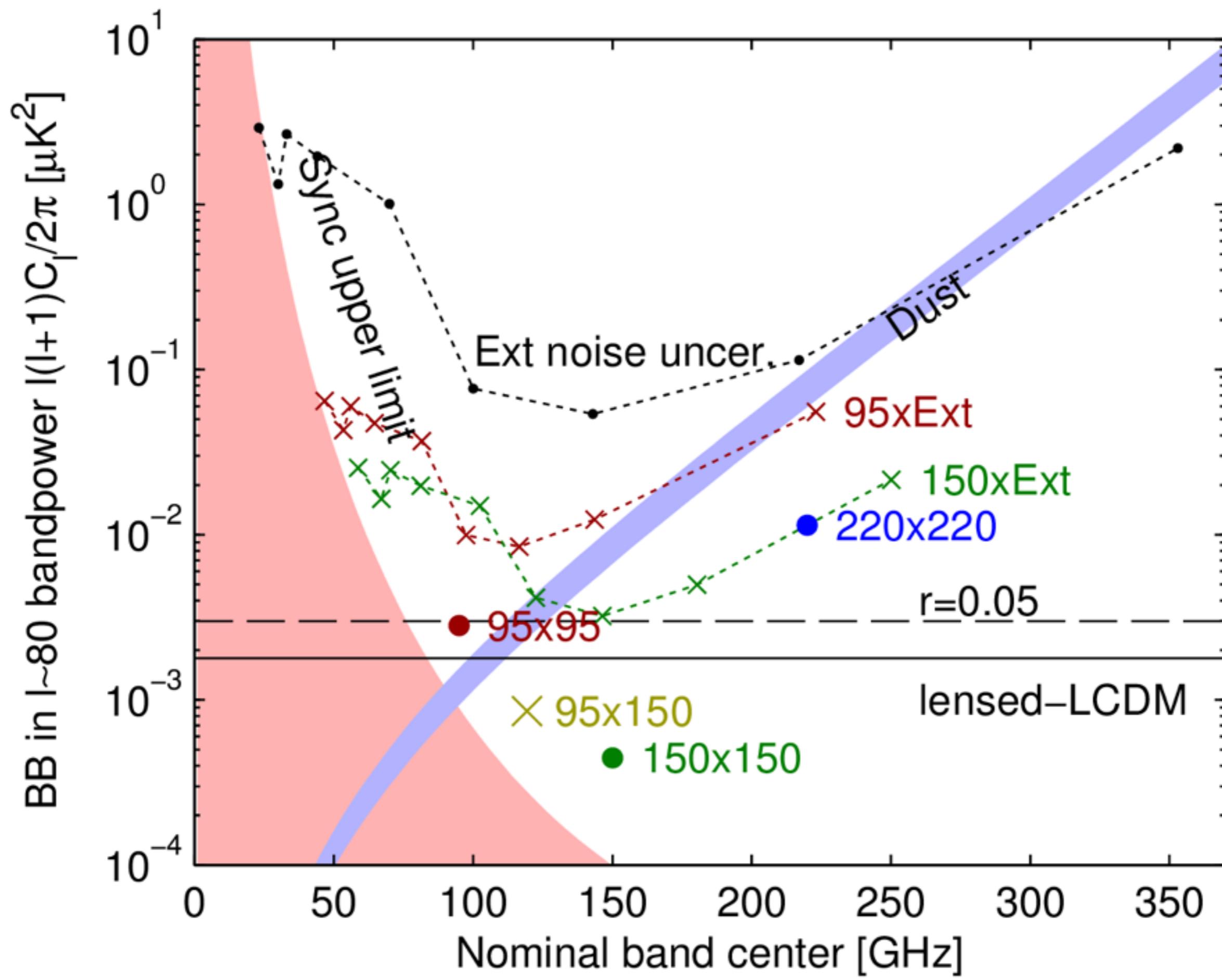


2015

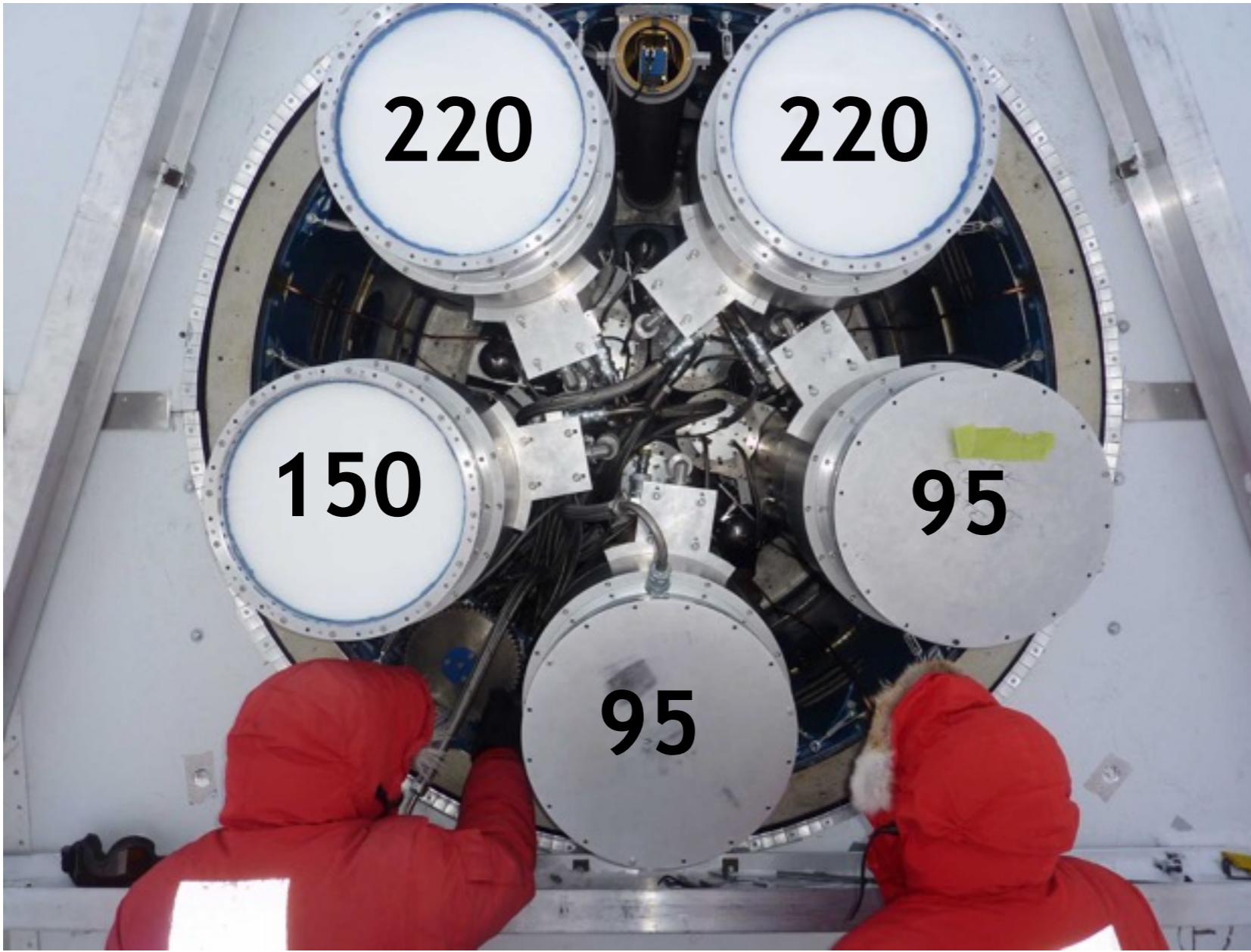
Ground windows





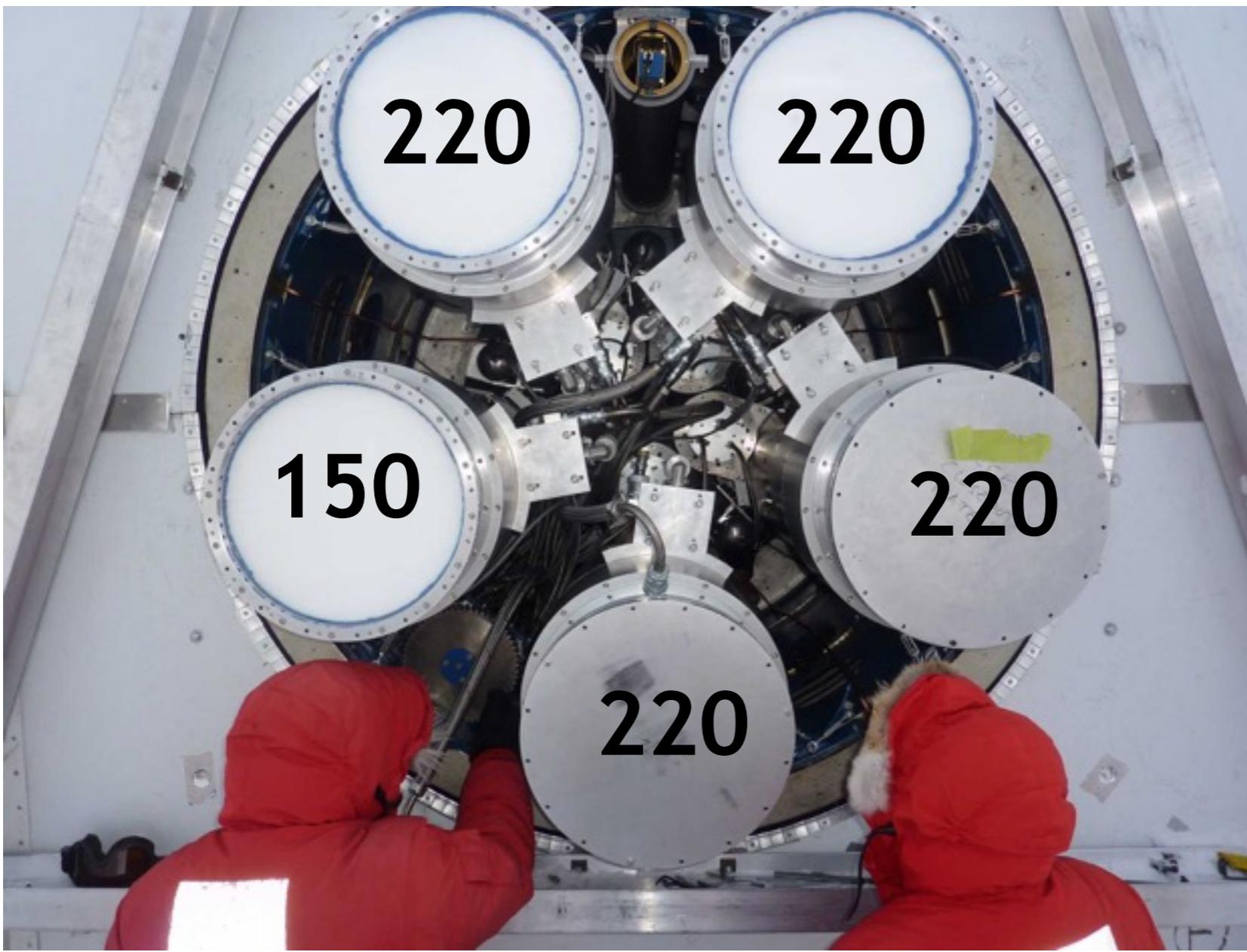


Keck Array Frequency Coverage



2015

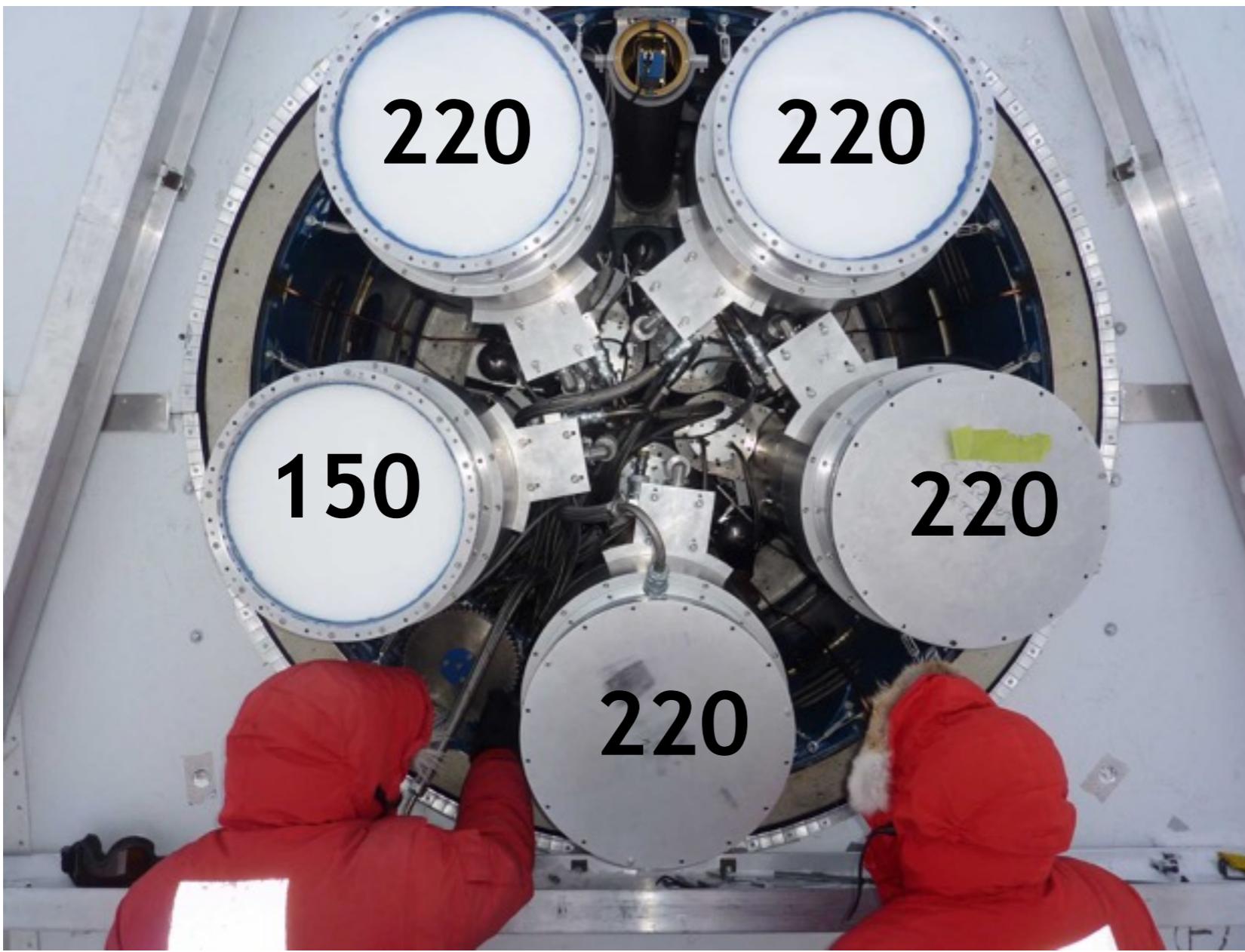
Keck Array Frequency Coverage



2016

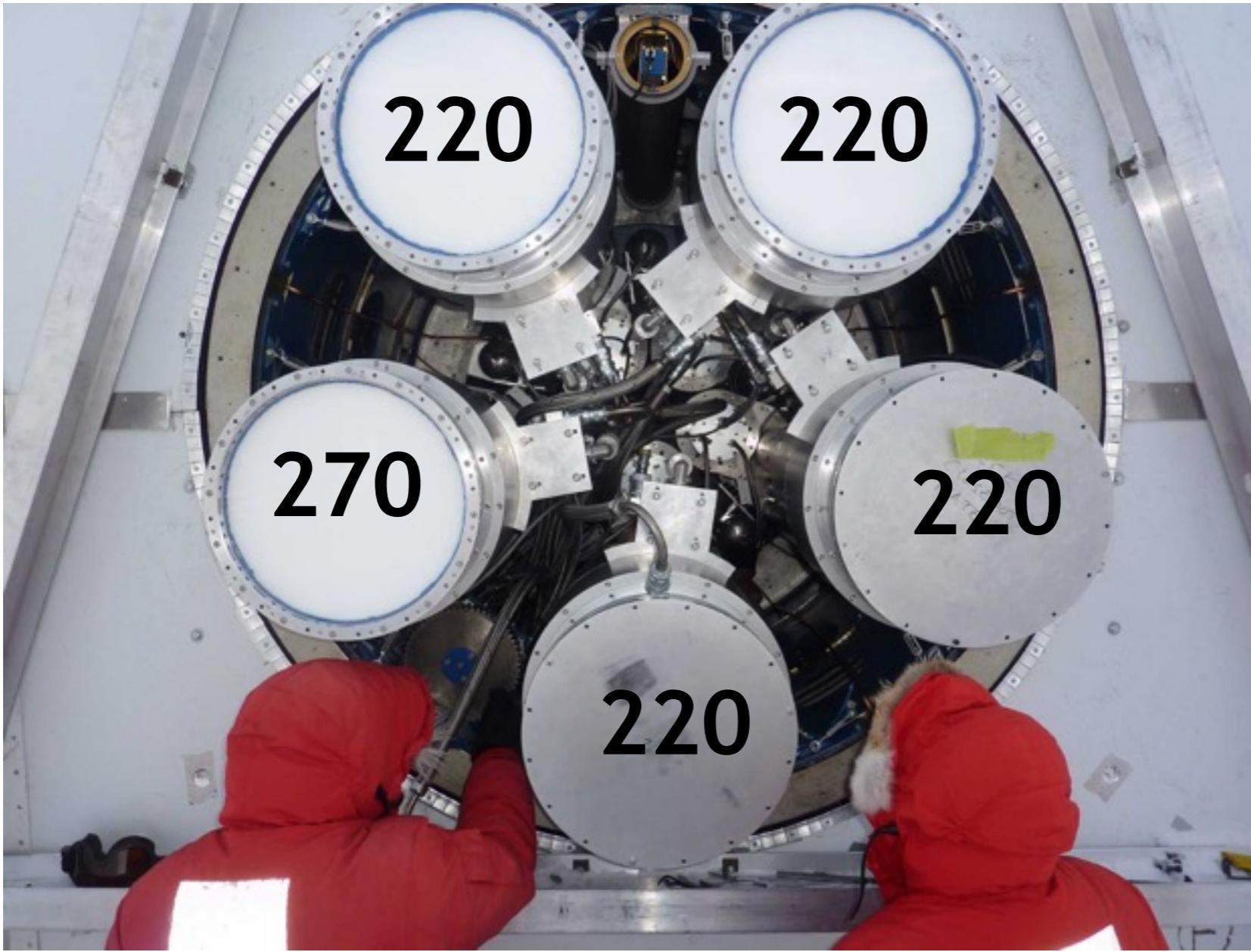
BICEP/Keck maps pager

Keck Array Frequency Coverage



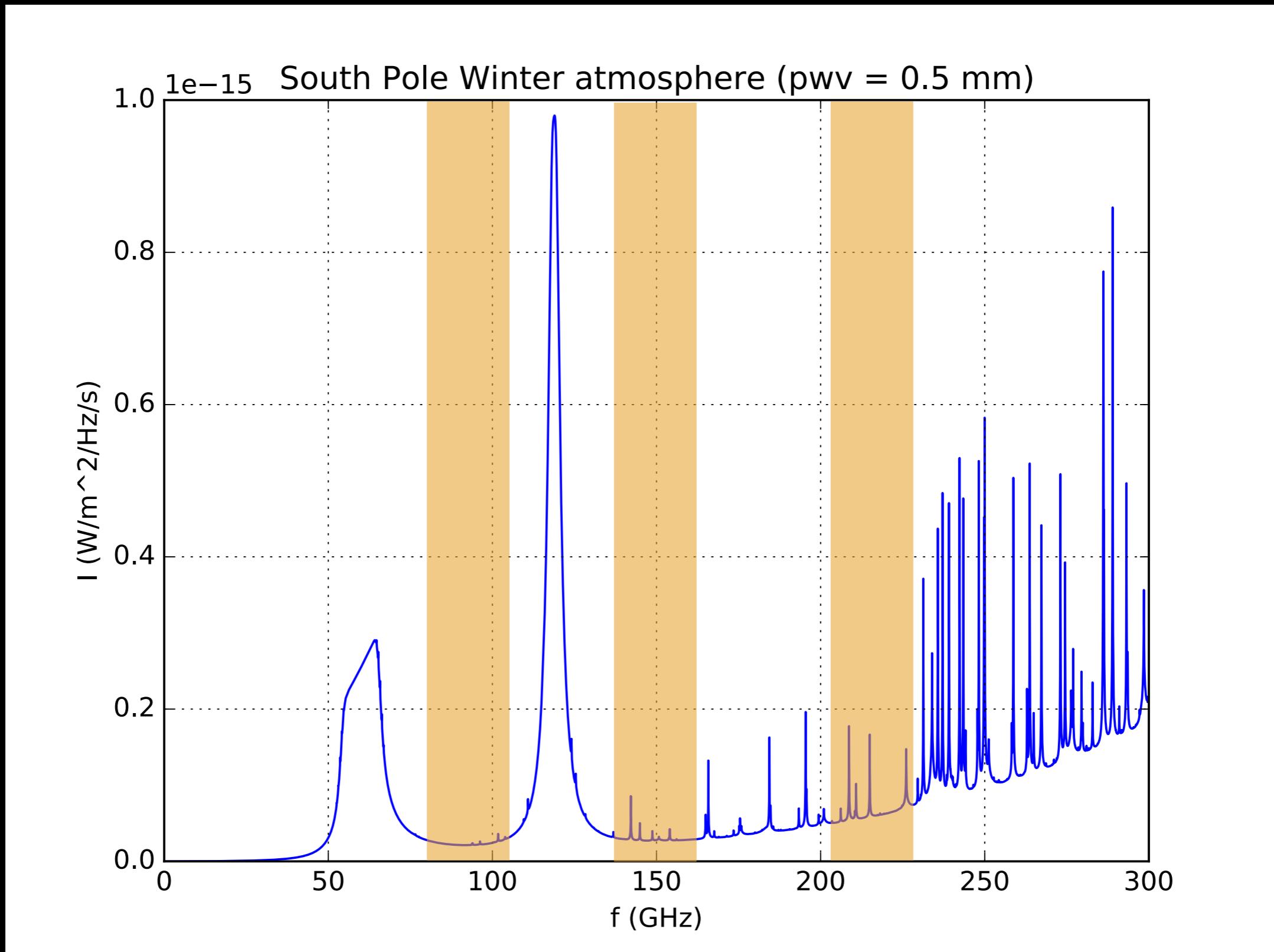
2016

Keck Array Frequency Coverage

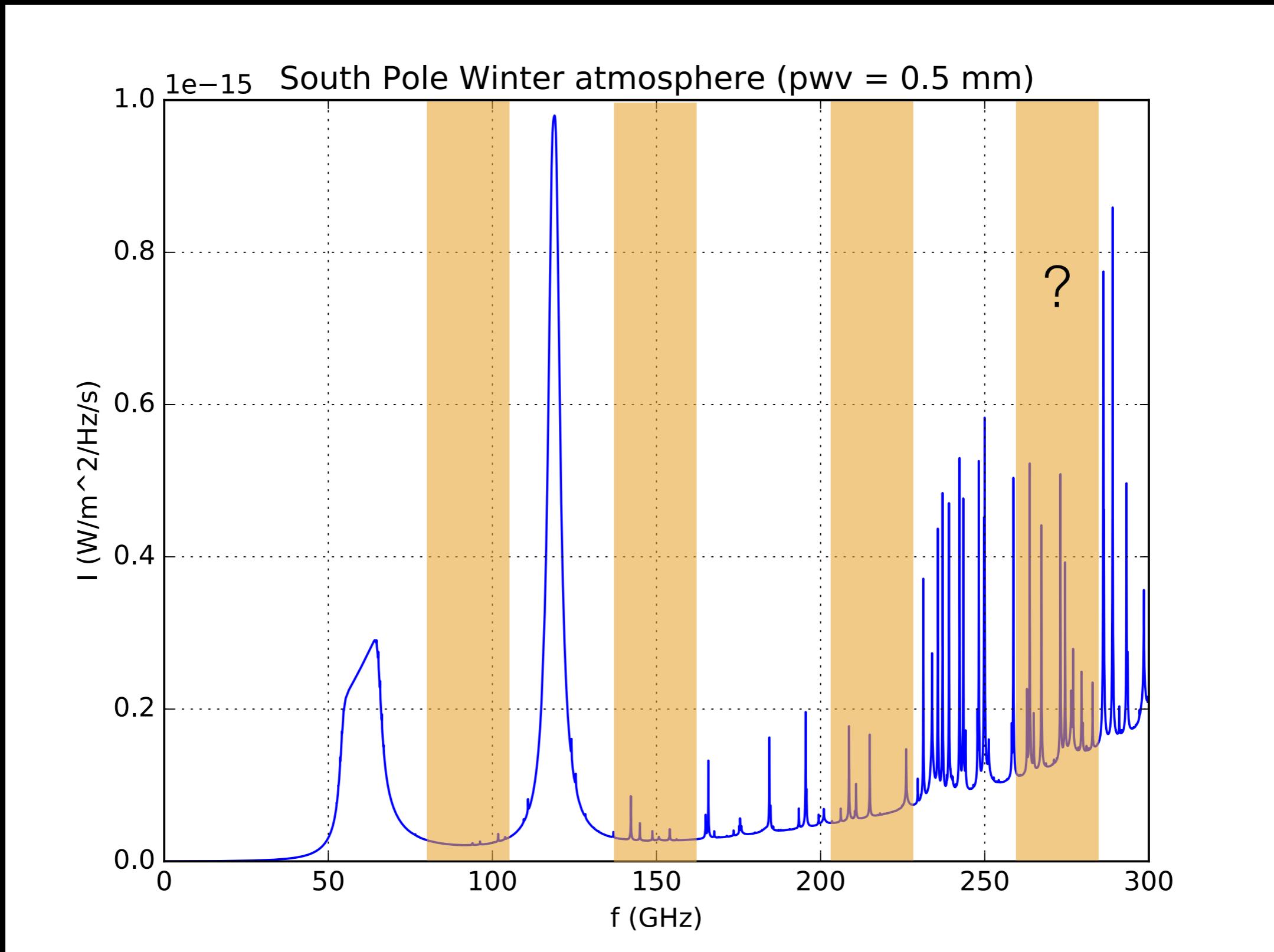


2017

Ground windows



Ground windows





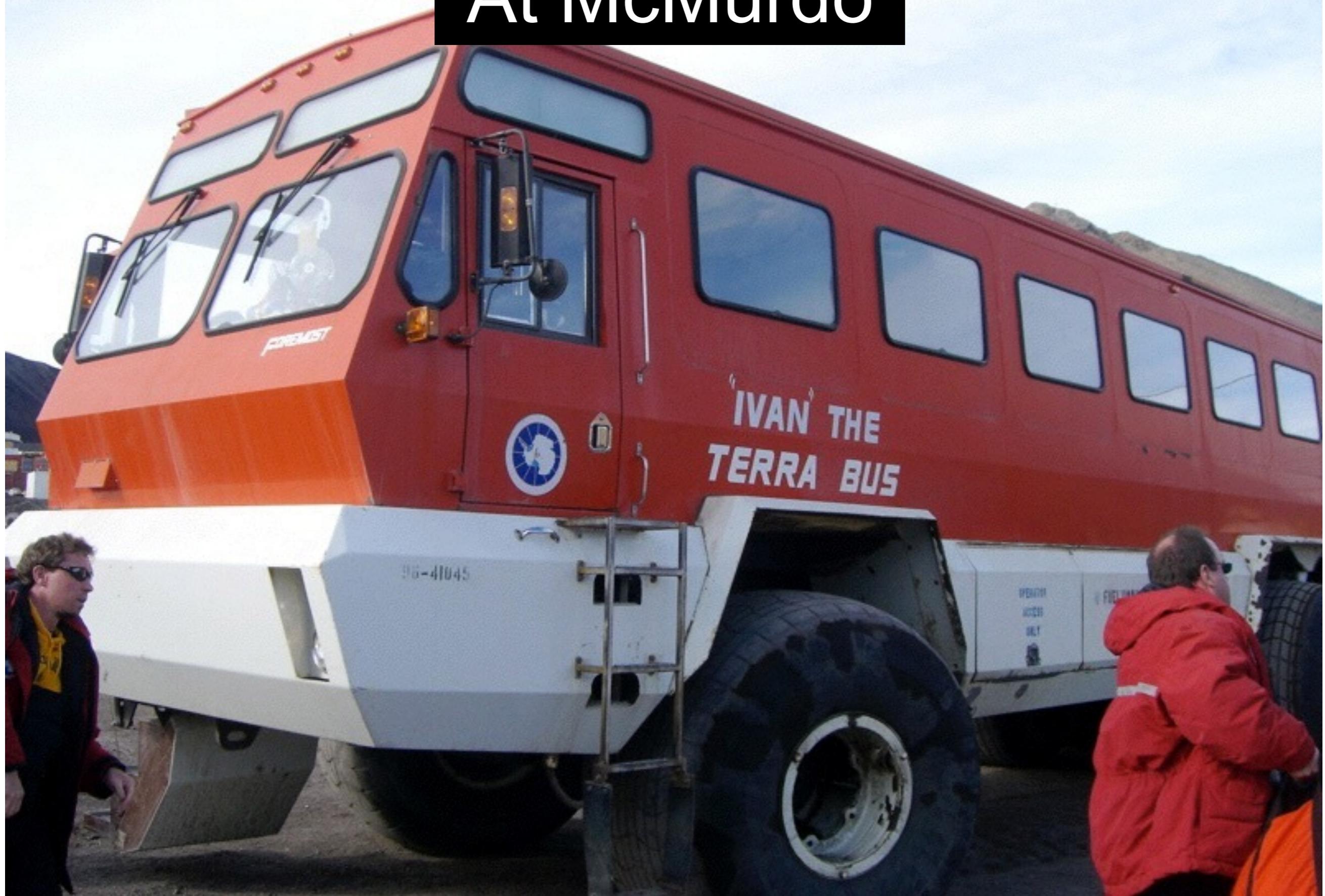
Inside C17



At McMurdo



At McMurdo



At McMurdo



At McMurdo



At McMurdo



At McMurdo



At McMurdo



LC-130 Herc



LC-130 Herc



Transantarctic Mountains



Transantarctic Mountains



Getting up on the plateau...



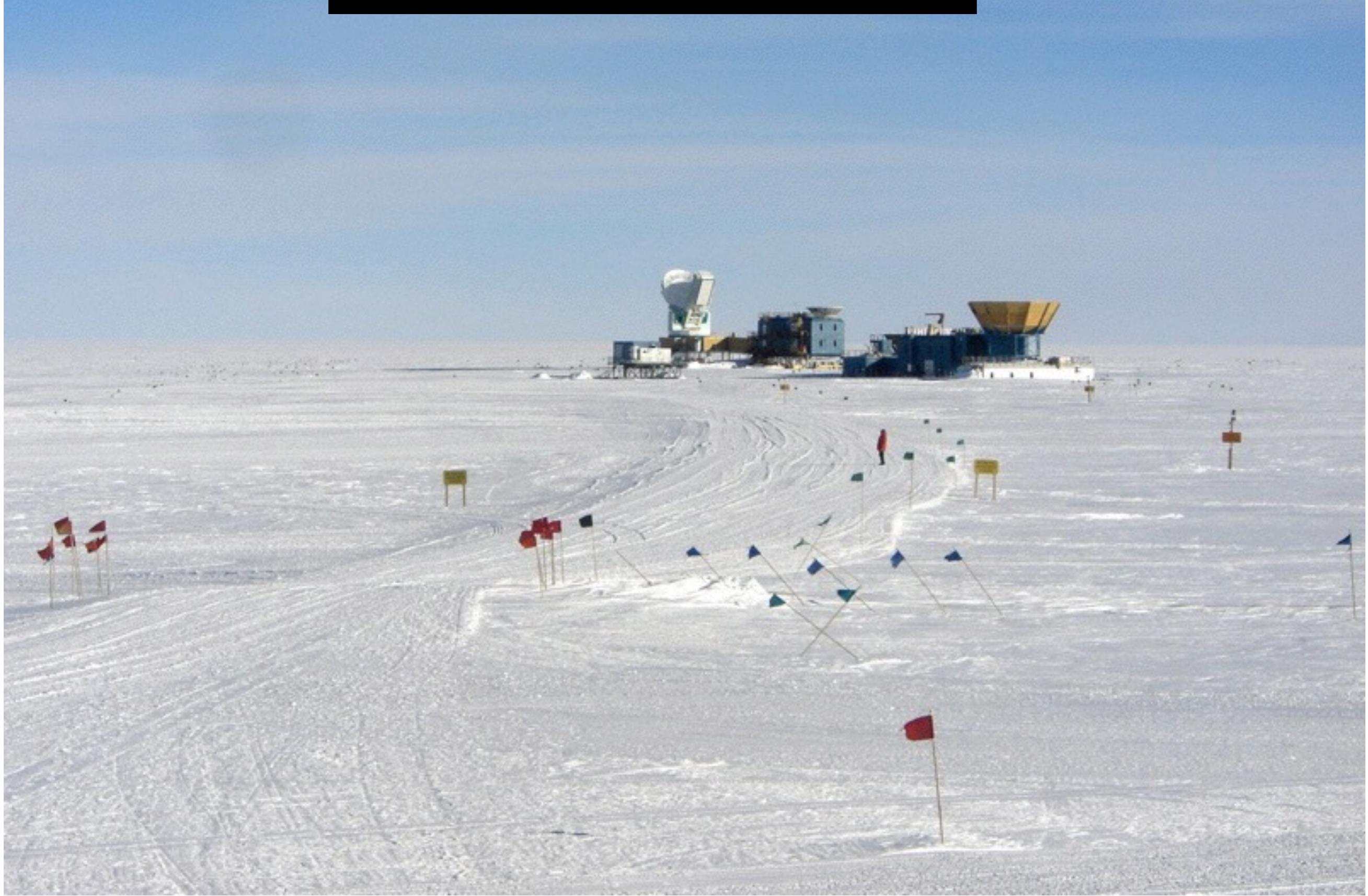
Last bit of land...



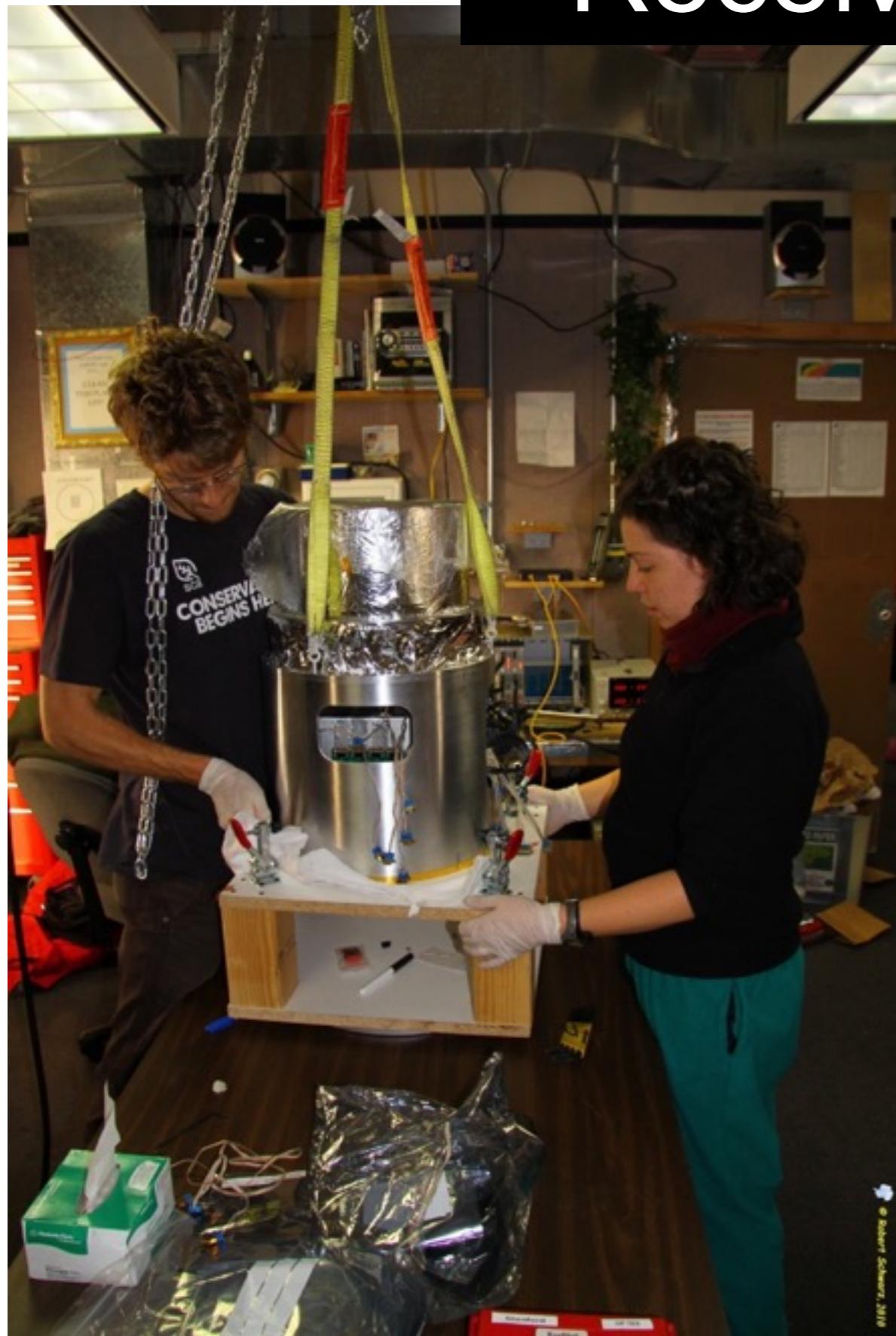
Arrival at Pole



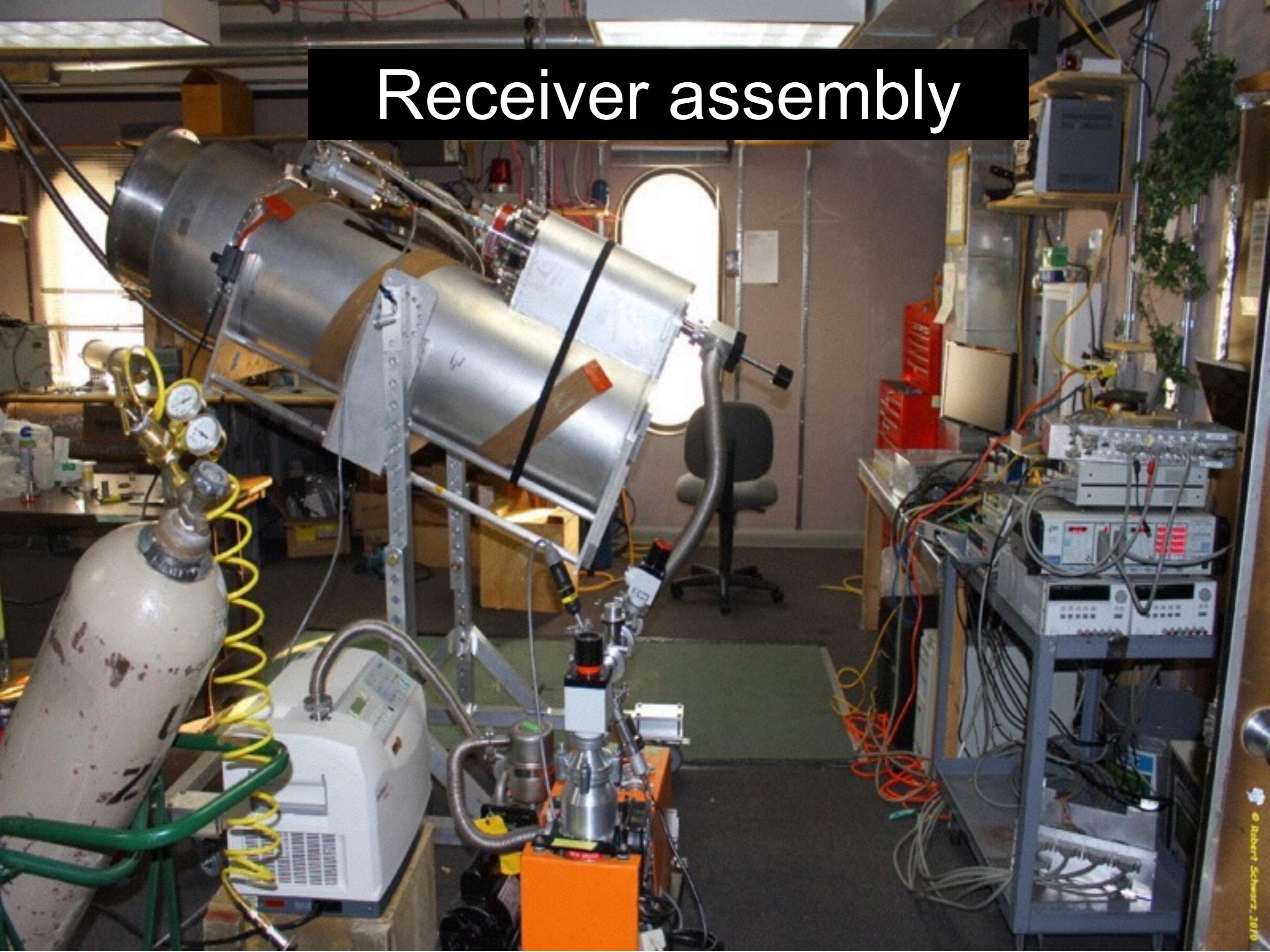
Dark Sector



Receiver assembly



Receiver assembly

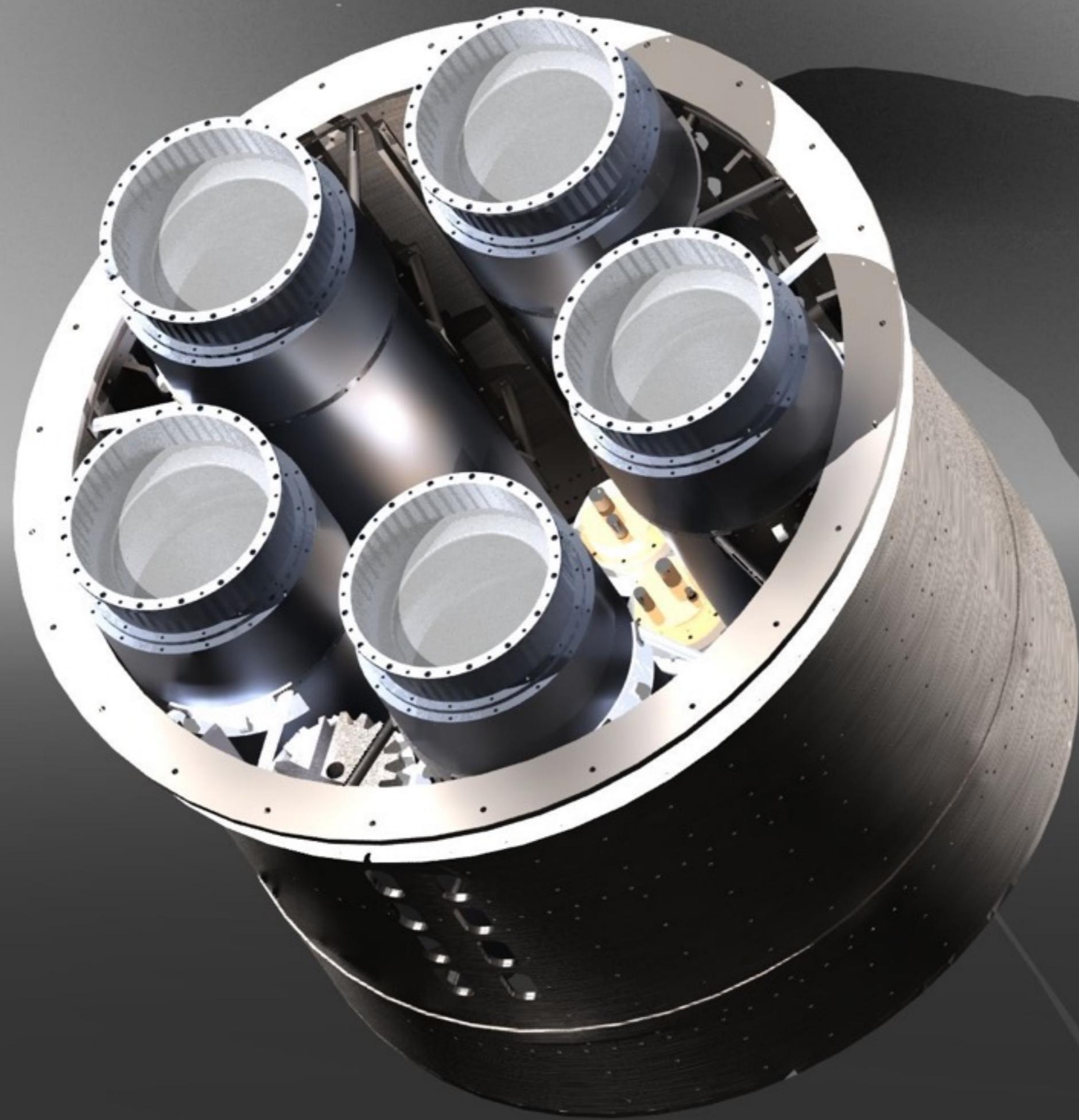


Receiver loading



Receiver loading





Me



CHRIS SHEEHY

BICEP3 Beam Calibration

Far field beam mapping:
microwave source mounted here

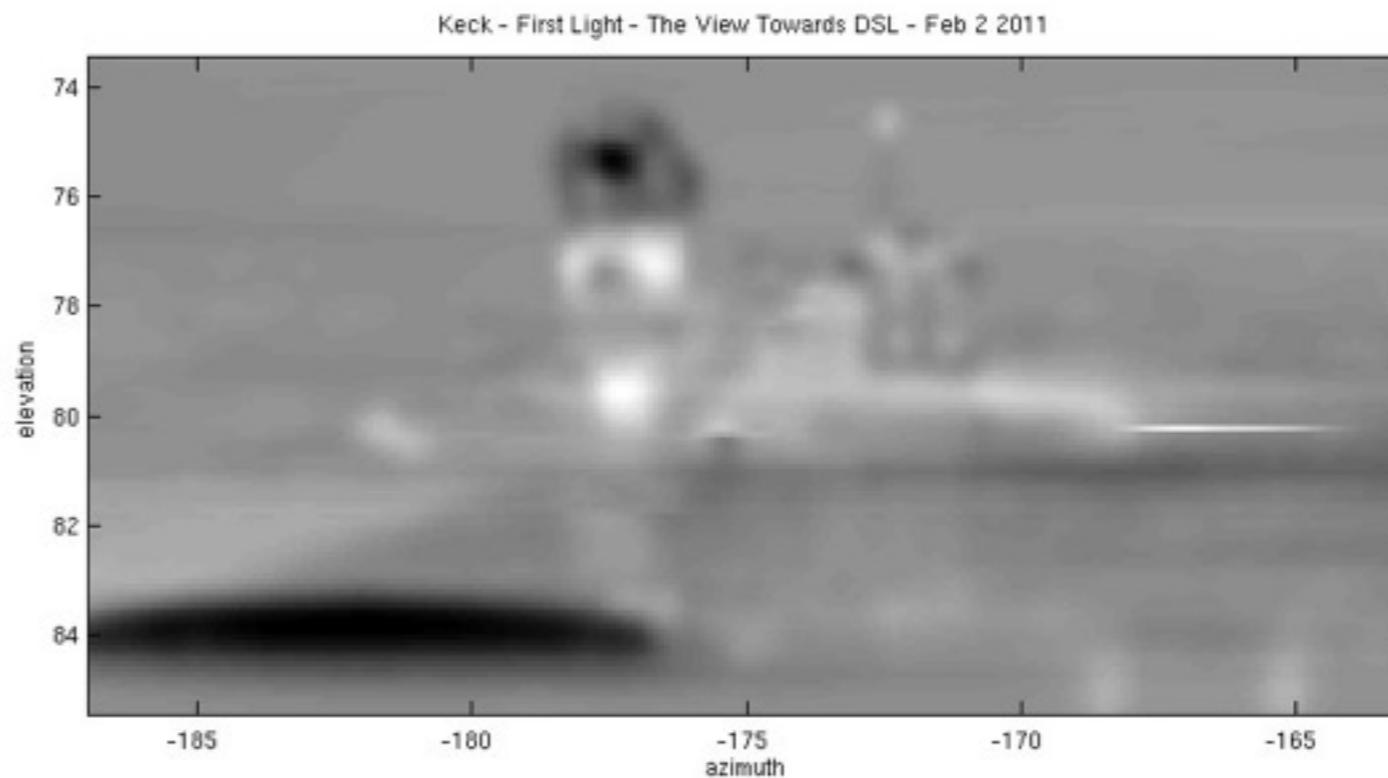


Keck Beam Calibration



Beam map raw data

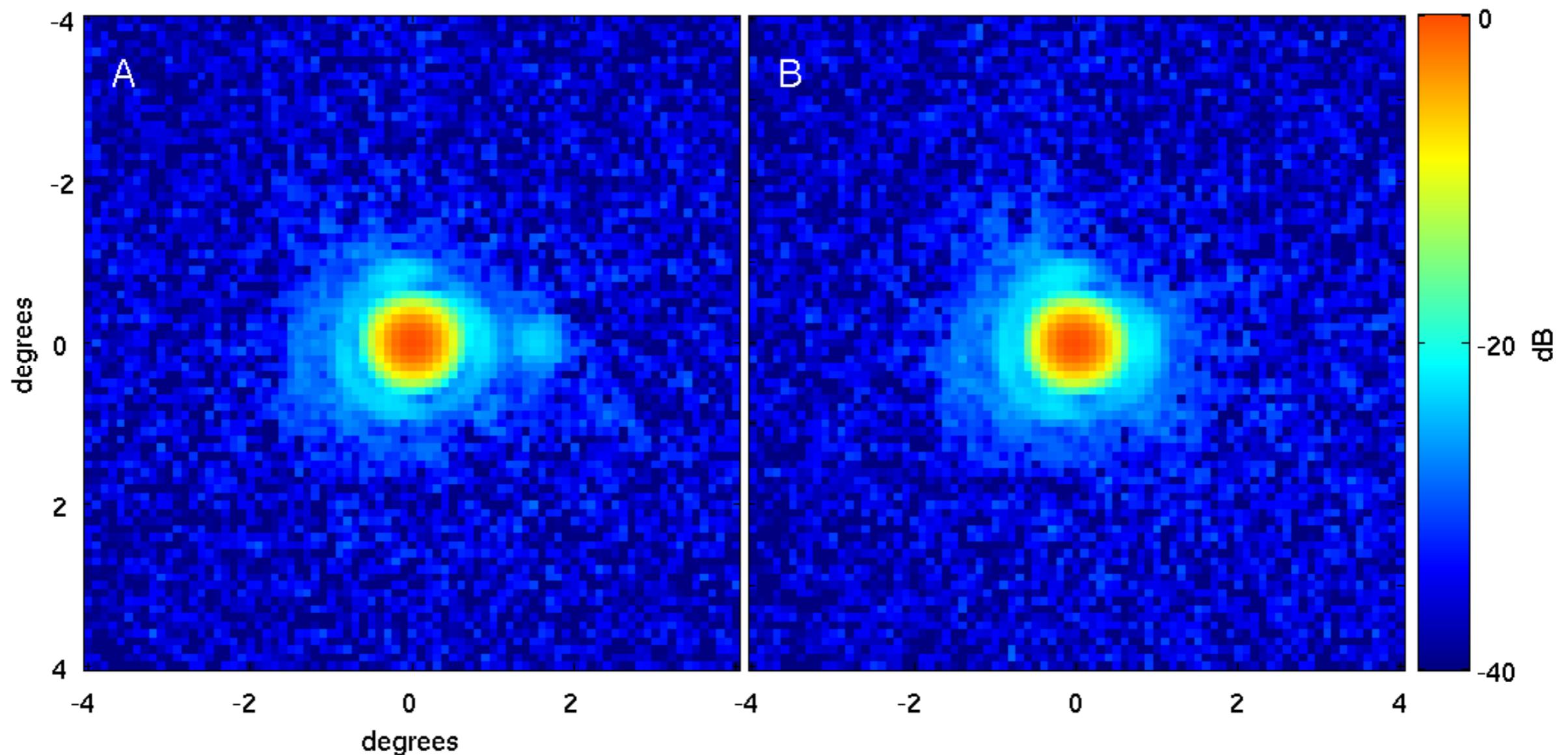
150 GHz



Visible light

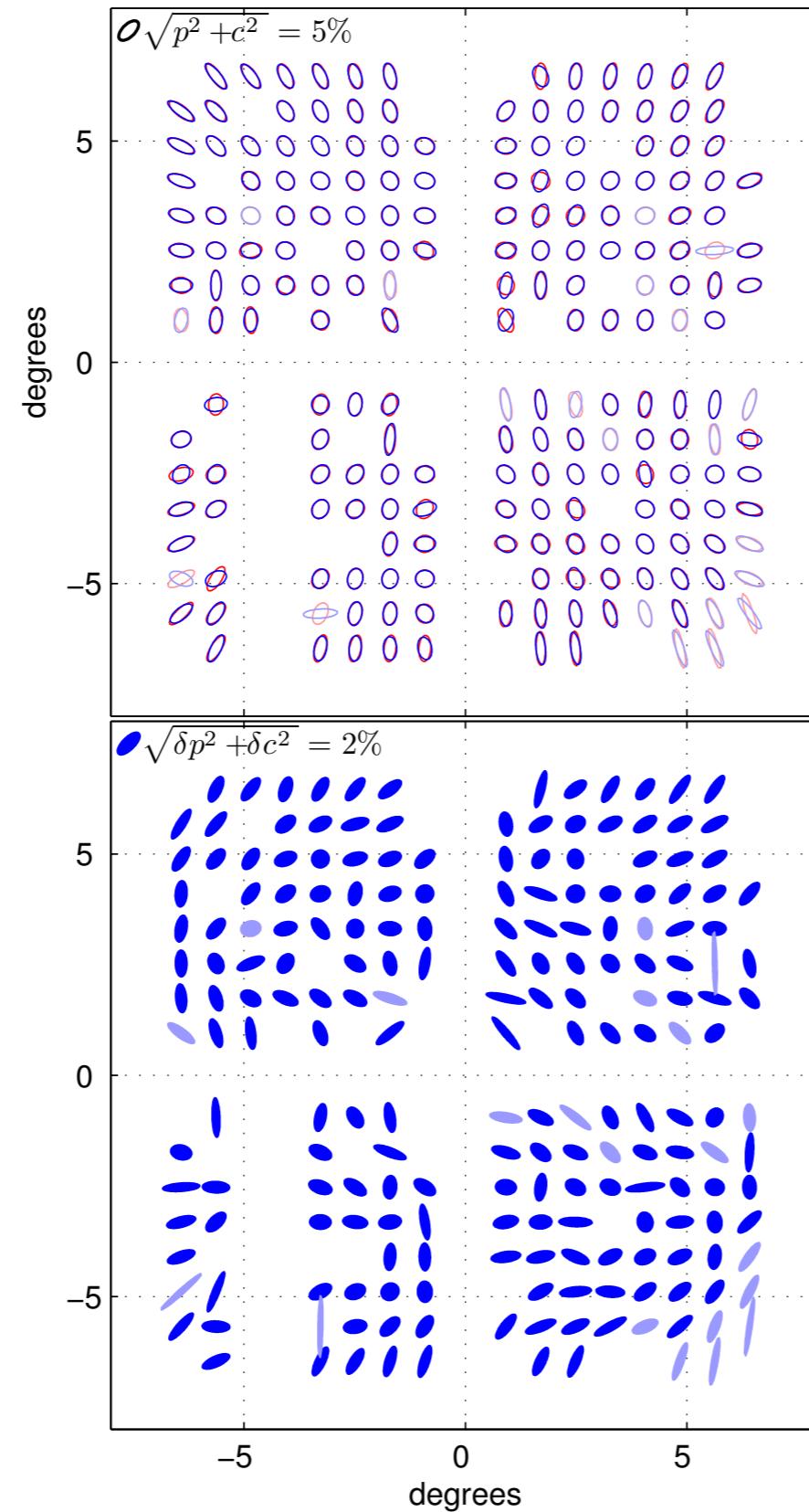


Beam maps



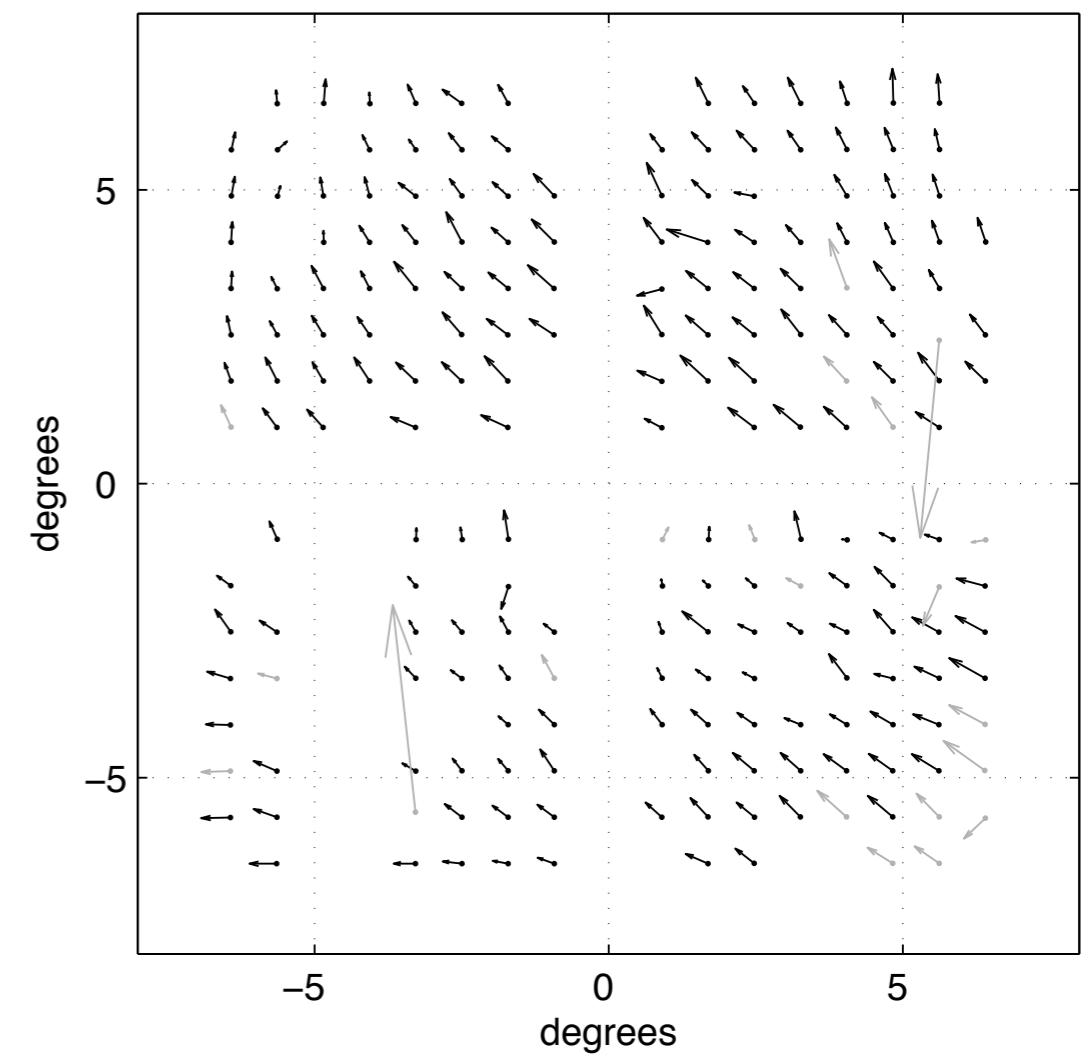
We know our beam shapes

beam ellipticity

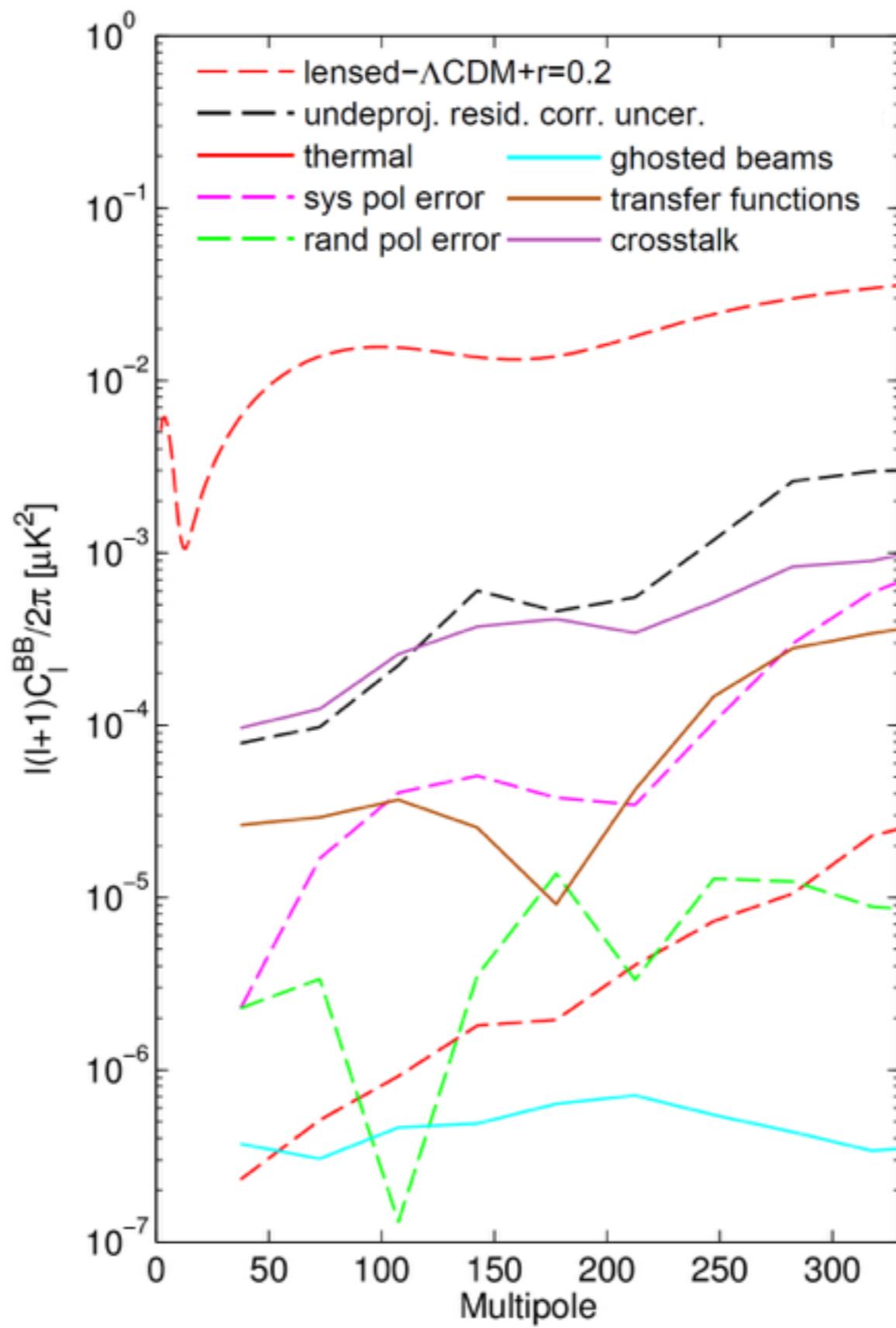


differential ellipticity

differential beam centroids (x20)



Systematics



Mount and motion control upgrades



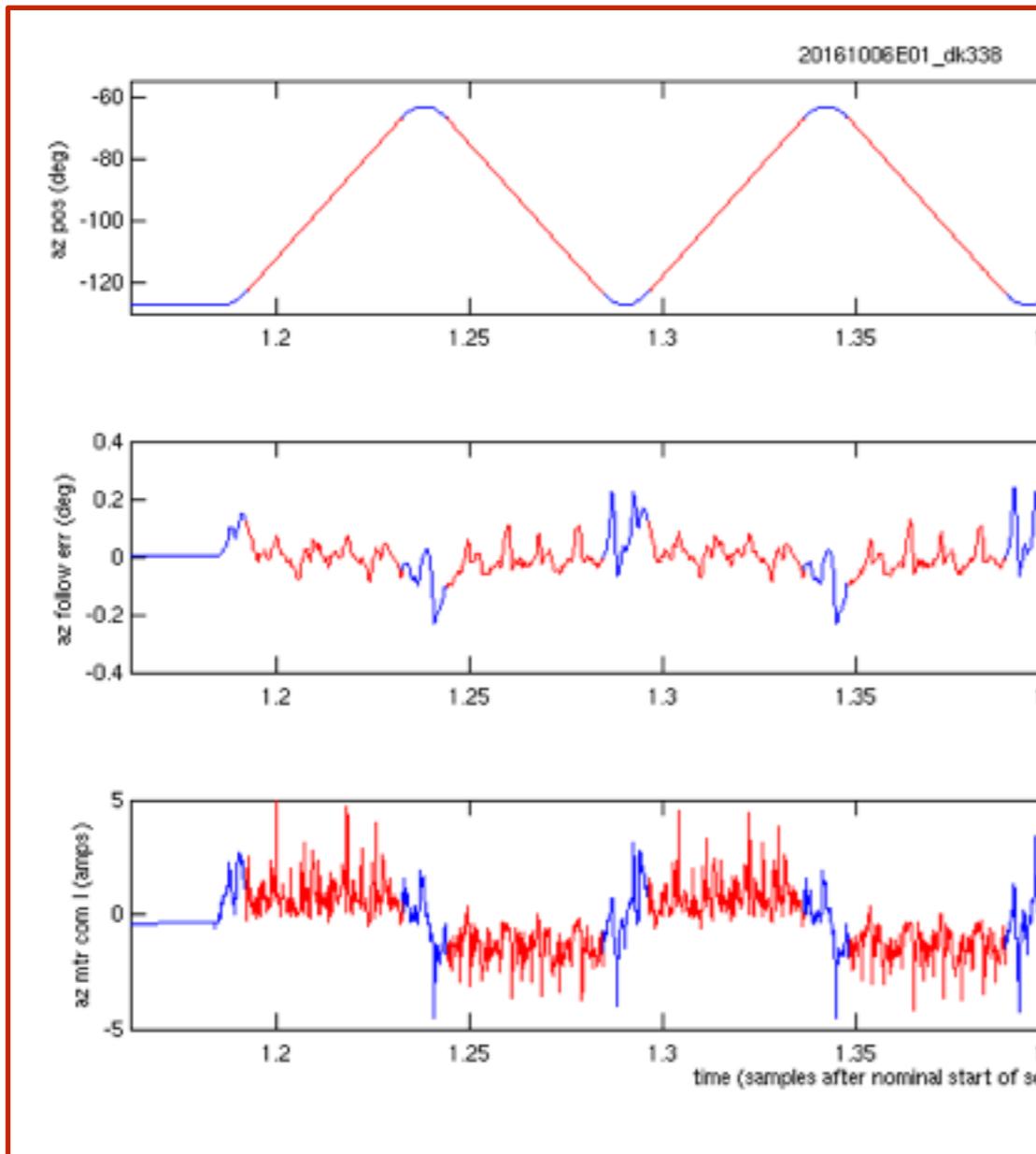
Mount and motion control upgrades



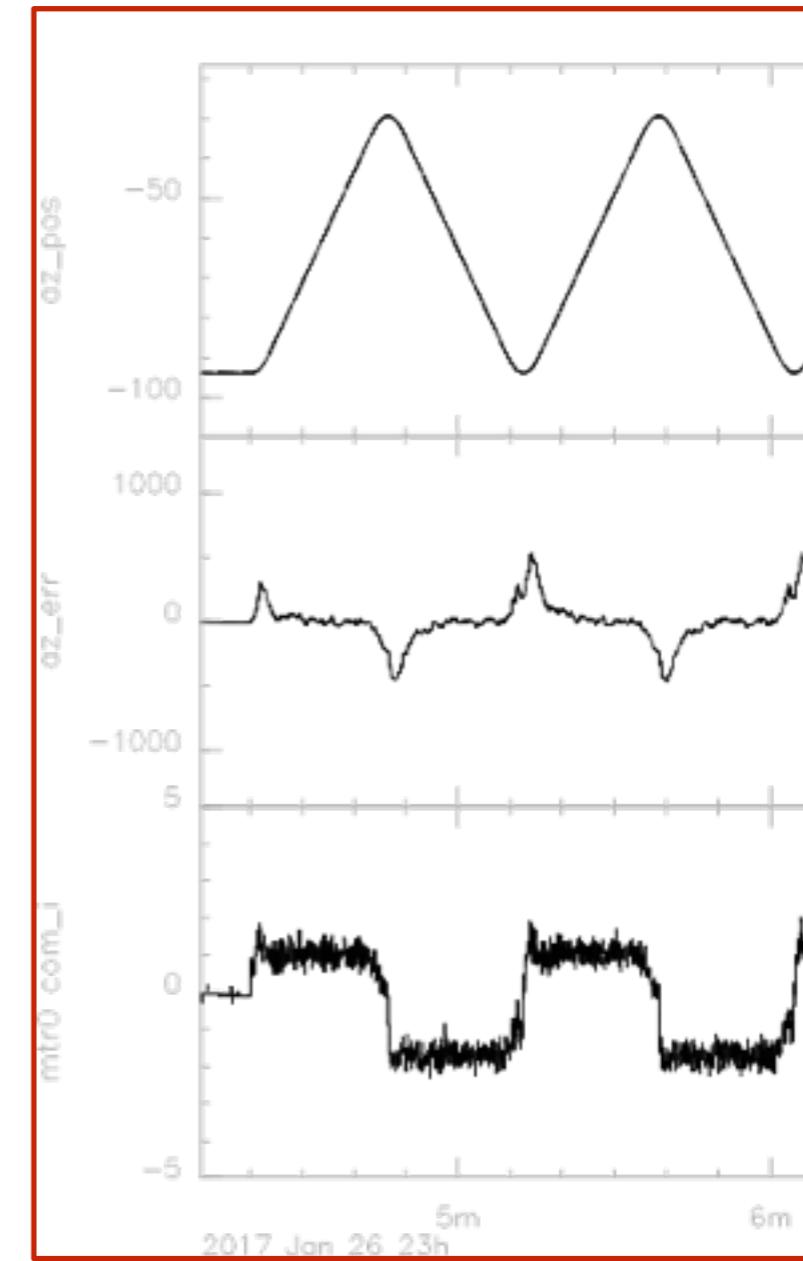
Mount and motion control upgrades



Mount and motion control upgrades



Before



After

Prep for BICEP Array, now funded

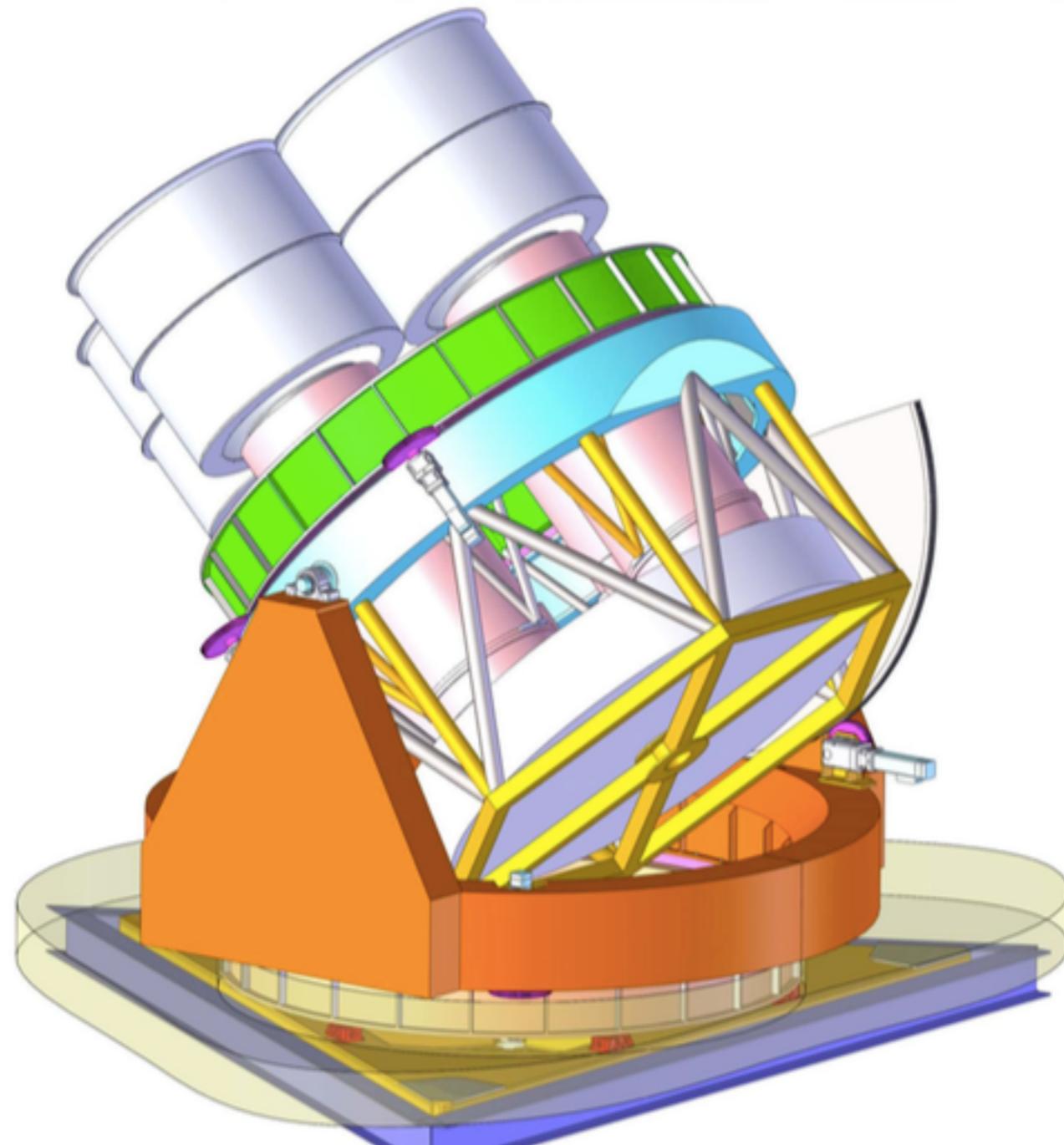


Figure 1: BA mount rear iso-view at 45° elevation



video: Robert Schwartz